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UNIVERSITÉ  
DE GENÈVE



# Host's star and habitability

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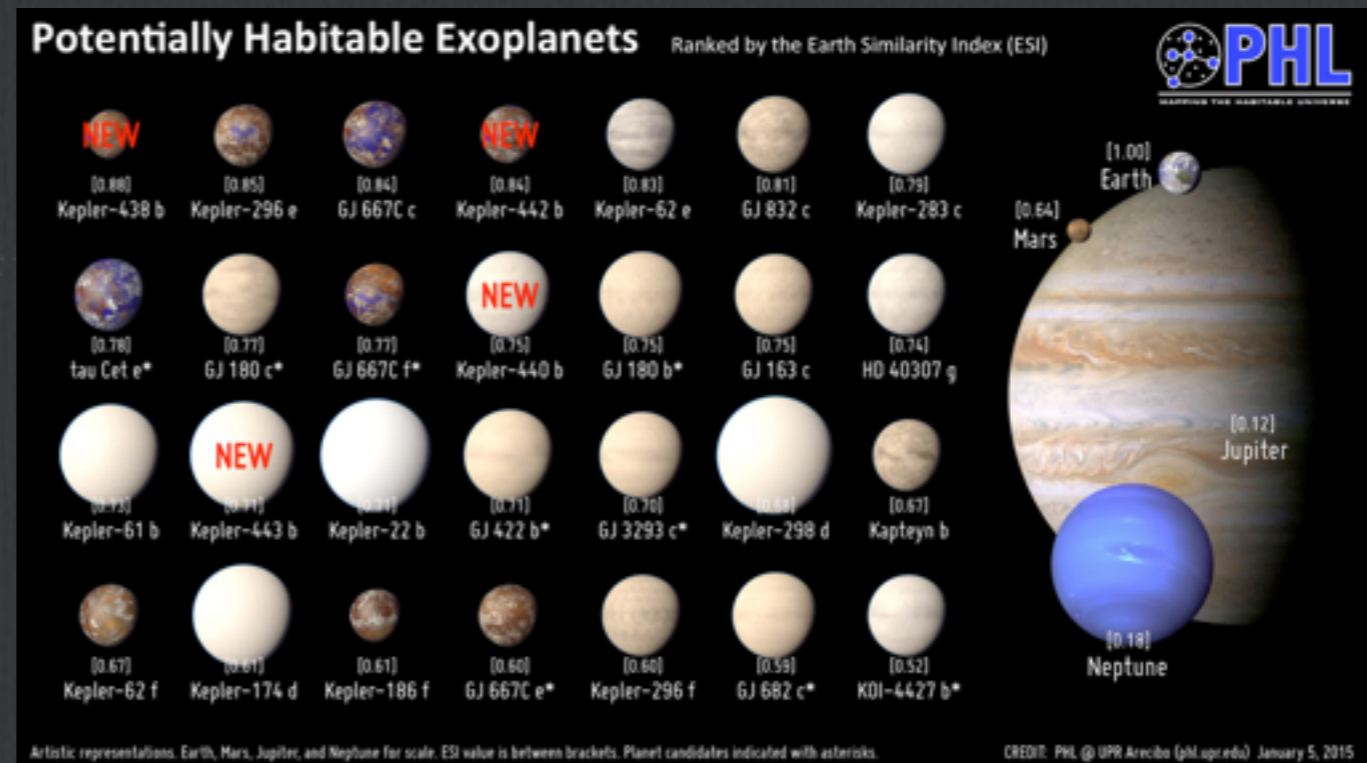
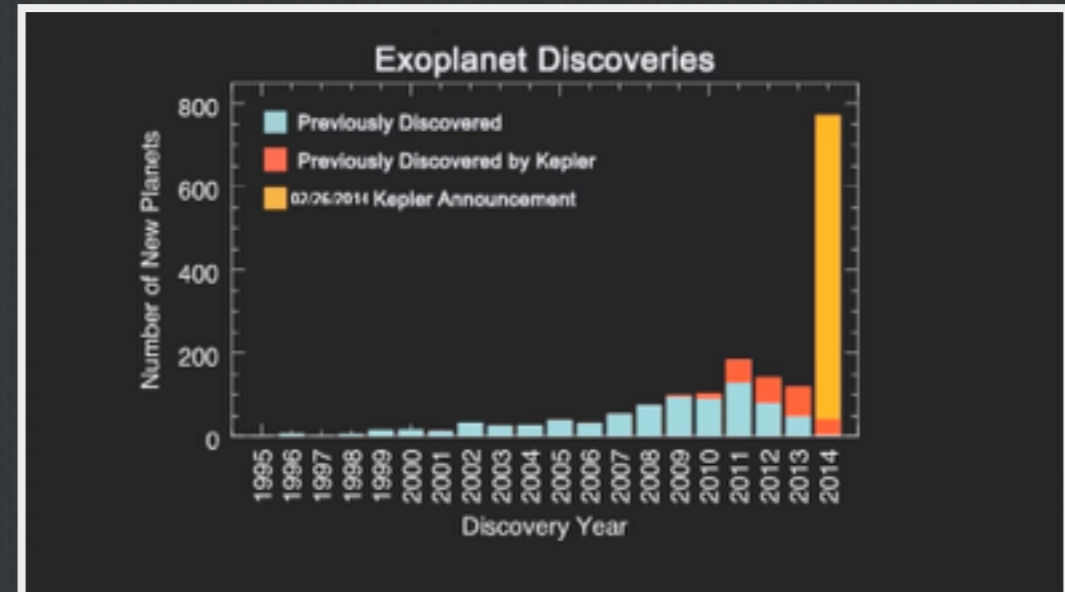
Ana Palacios (LUPM), Louis Amard (LUPM/Geneva), Sacha Brun (AIM Paris-Saclay), Stephane Mathis (AIM Paris-Saclay)

# brief Context

☐ 1642 planets confirmed

☐ ~30 in the habitable zone

☐ => need to understand these objects and their behavior in such interesting regions



Ref : [exoplanets.org](http://exoplanets.org)

# Literature

Not exhaustif

## Linsenmeier et al. 2015

- « Commonly, a planet is considered habitable if its surface conditions allow for the existence of liquid water. »

## Torres et al. 2015

- « Public curiosity and scientific interest have motivated efforts in the last few years to find and confirm rocky planets similar in size to the Earth that are orbiting in the so-called habitable zone (HZ) of their parent stars, usually taken in this context to be the region in which water on the surface can be in a liquid state. »

## Bolmont et al. 2014 (SF2A)

- « We define here the habitable zone (HZ) as the region around a star where a planet with the right atmosphere can potentially sustain surface liquid water (Kasting et al. 1993; Selsis et al. 2007). »

## Kasting et al. 2013/2014

- « The habitable zone (HZ) around a star is typically defined as the region where a rocky planet can maintain liquid water on its surface. That definition is appropriate, because this allows for the possibility that carbon-based, photosynthetic life exists on the planet in sufficient abundance to modify the planet's atmosphere in a way that might be remotely detected »

**The region where a rocky planet can maintain liquid water on its surface.**

# Habitability

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Habitable zone + « life's ingredients »

- Habitable zone =  $T_{\text{eff}}/L_{*} + S_{\text{eff}}$  (given planet atmosphere)
  - stellar point of view
  
- « ingredients » = tidal interaction, magnetic field, volcanic activity ...
  - ingredients required to host « life »
  - planetary point of view

# Habitable zone...



... = snapshot?

- No!
  - follow intrinsic evolution of star
  - previous evolution?
- Need stellar models!**
- Require precise
  - mass
  - metallicity
  - age

...sorry

# Reference grid

## STAREVOL

$$\alpha_c = 1.7020$$

$$Y = 0.2689$$

$$Z = 0.0134$$

Non-grey atmosphere

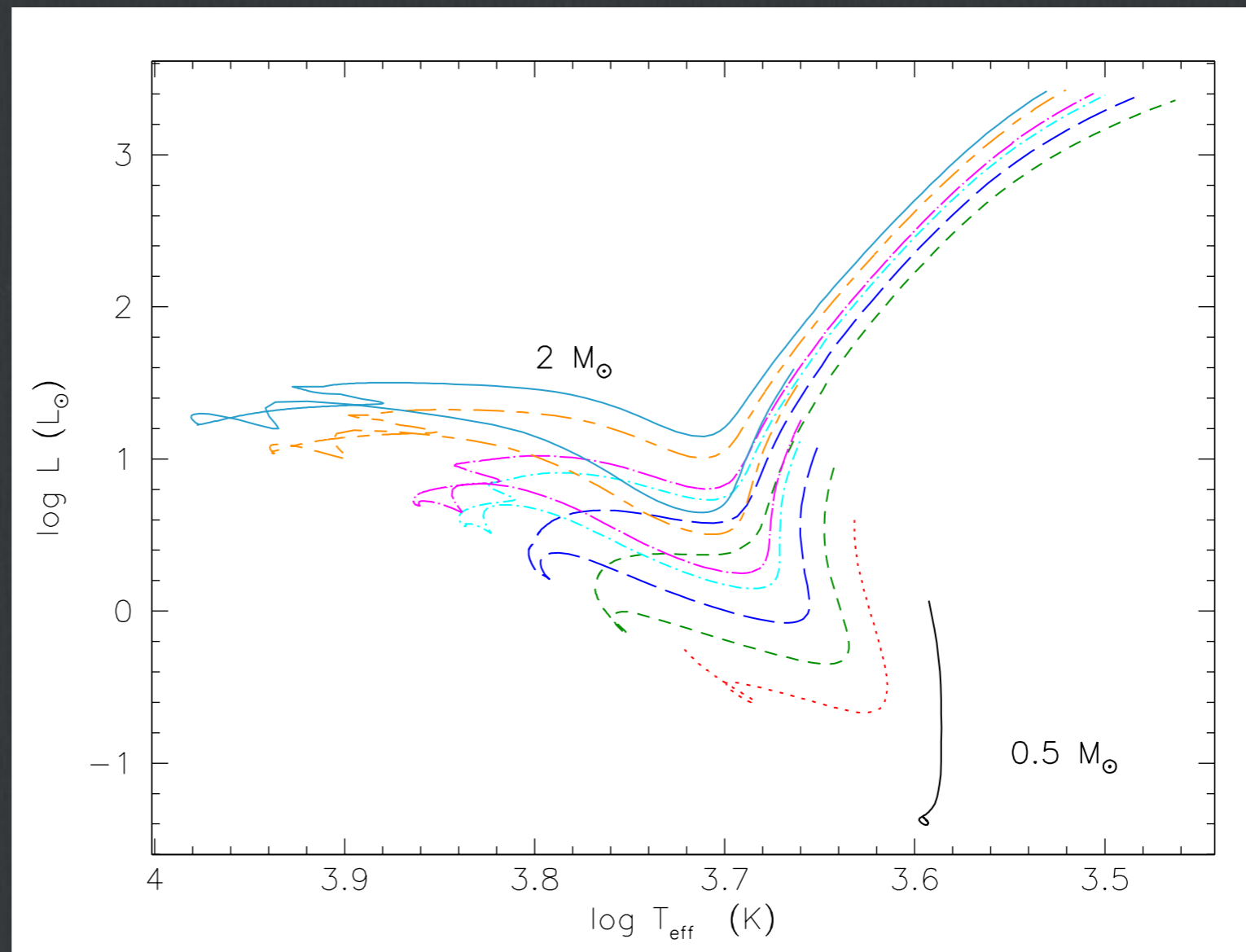
Internal transport of AM

Rotation

AM and mass loss

Impact of centrifugal effect  
on effective gravity

...



$$0.5 < M_*/M_\odot < 2$$

See Lagarde et al. (2012) and Amard et al. (submitted) for more details

# Habitable Zone

Kopparapu et al. (2013)

$$S_{eff} = S_{eff\odot} + aT_* + bT_*^2 + cT_*^3 + \dots$$

$$2600 \text{ K} \leq T_{eff} \leq 7200 \text{ K}$$

$S_{eff\odot}, a, b, c = f(\text{planet atmosphere})$

$$S_{eff} = \frac{F_{IR}}{F_{inc}}$$

$$d = \left( \frac{L / L_{\odot}}{S_{eff}} \right)^{0.5} \text{ AU}$$

Kasting et al. (1993)

$S_{eff}$  = effective stellar flux =  $F_{out}/F_{in}$

→ ratio between the outgoing IR flux from the planet and the net incident stellar flux

To compute the effective stellar flux, a model for the outgoing IR flux of the planet is needed

→ 1-D radiative-convective climate models

## Habitable region limits

- $R_{in}$  = Runaway greenhouse : net positive feedback of GH effect ( $T_{surf} > 647 \text{ K}$ ), ocean evaporate entirely
- $R_{out}$  = Maximum greenhouse : Rayleigh scattering by  $\text{CO}_2$  reduce GH ( $T_{surf} = 273 \text{ K}$ )

# Habitable zone

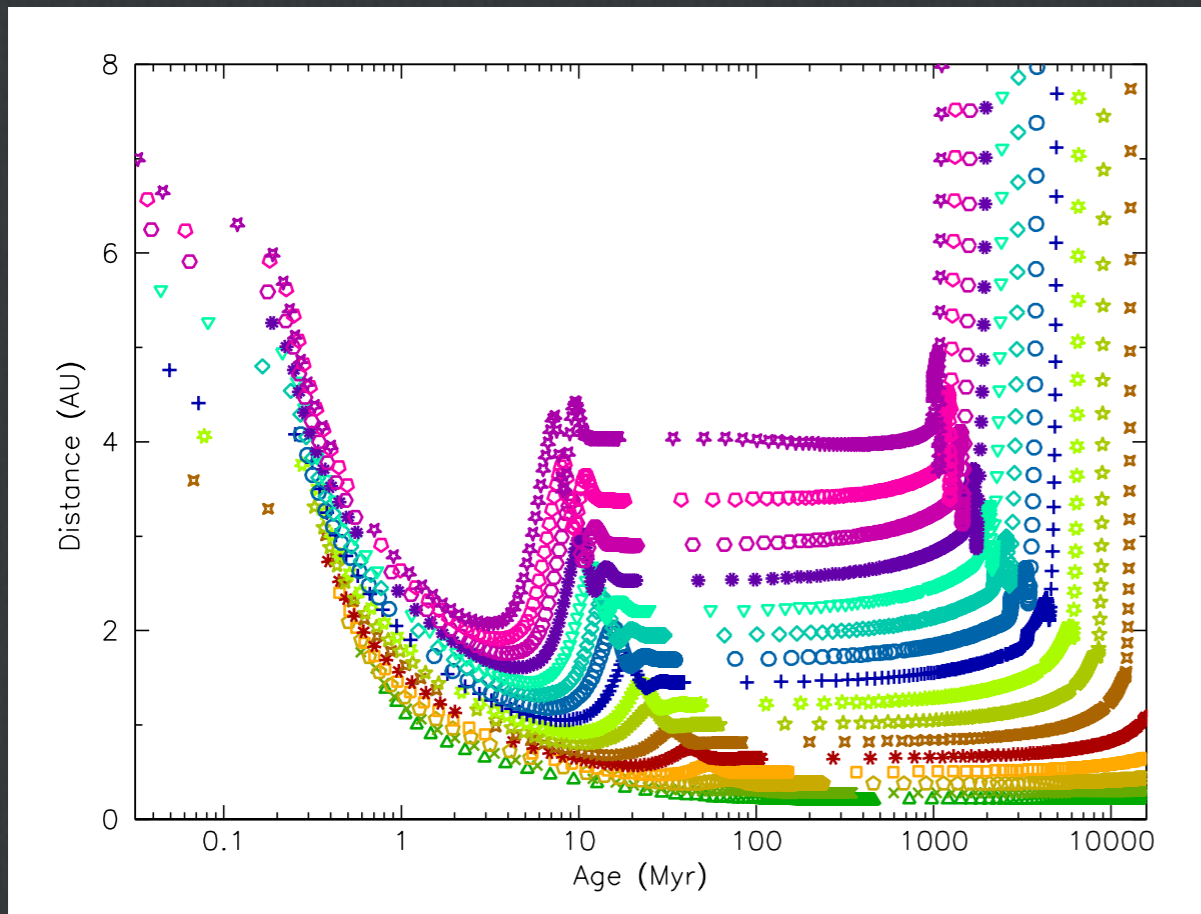
Impact of mass

0.5-2  $M_{\odot}$

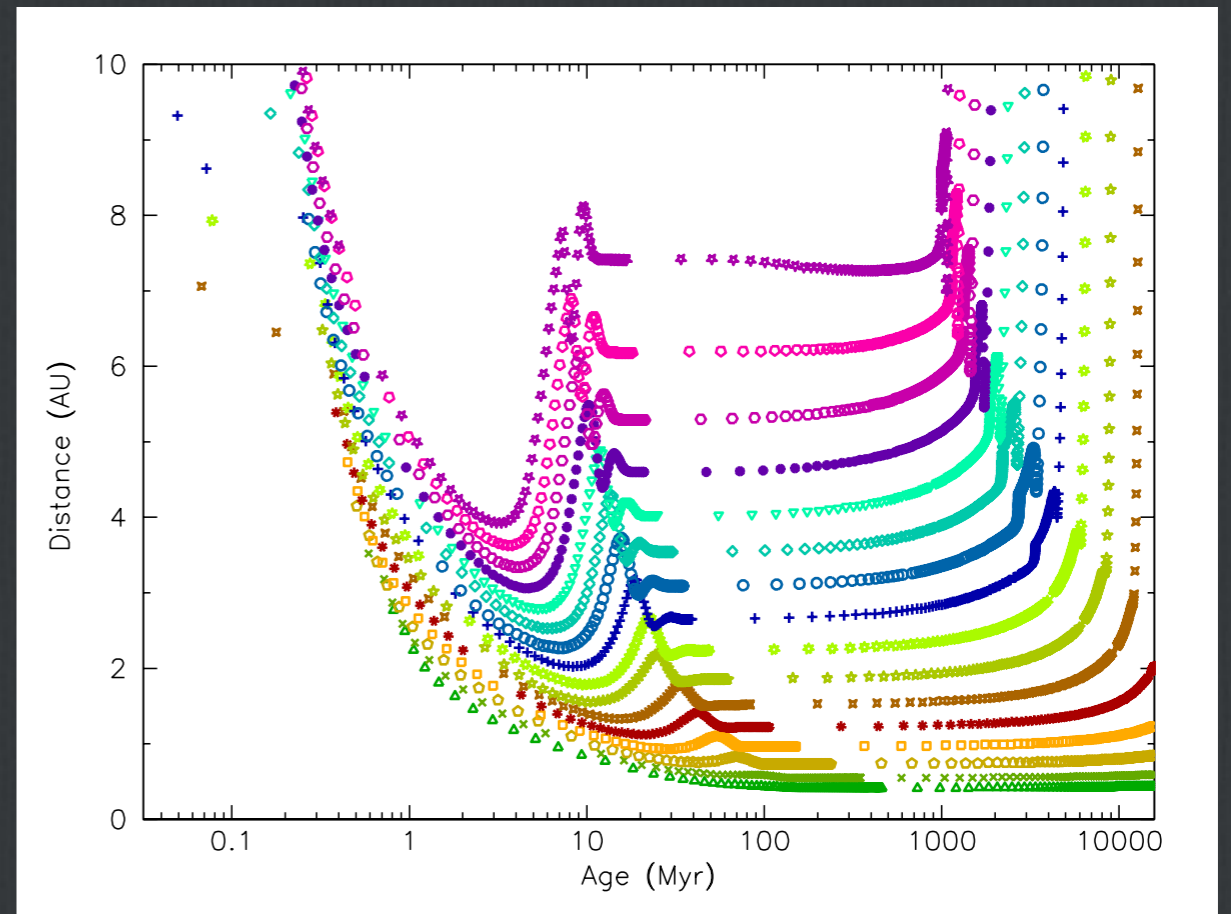
$Z_{\odot}$

No rotation

1  $M_{\oplus}$



$HZ_{in}$



$HZ_{out}$

- **Increases** towards higher masses
- **Shape** depends on the stellar mass
- **Correct** estimation of mass = **crucial**

Table 1. Size of the HZ as a function of stellar mass.

$\Delta HZ$	0.5 $M_{\odot}$	1 $M_{\odot}$	1.5 $M_{\odot}$	2 $M_{\odot}$
$\Delta HZ_{mean}$ (AU)	0.27	0.86	2.05	3.25
$\Delta HZ_{min}$ (AU)	0.2	0.65	1.2	1.85
$\Delta HZ_{max}$ (AU)	1.39	3.46	5.9	6.63



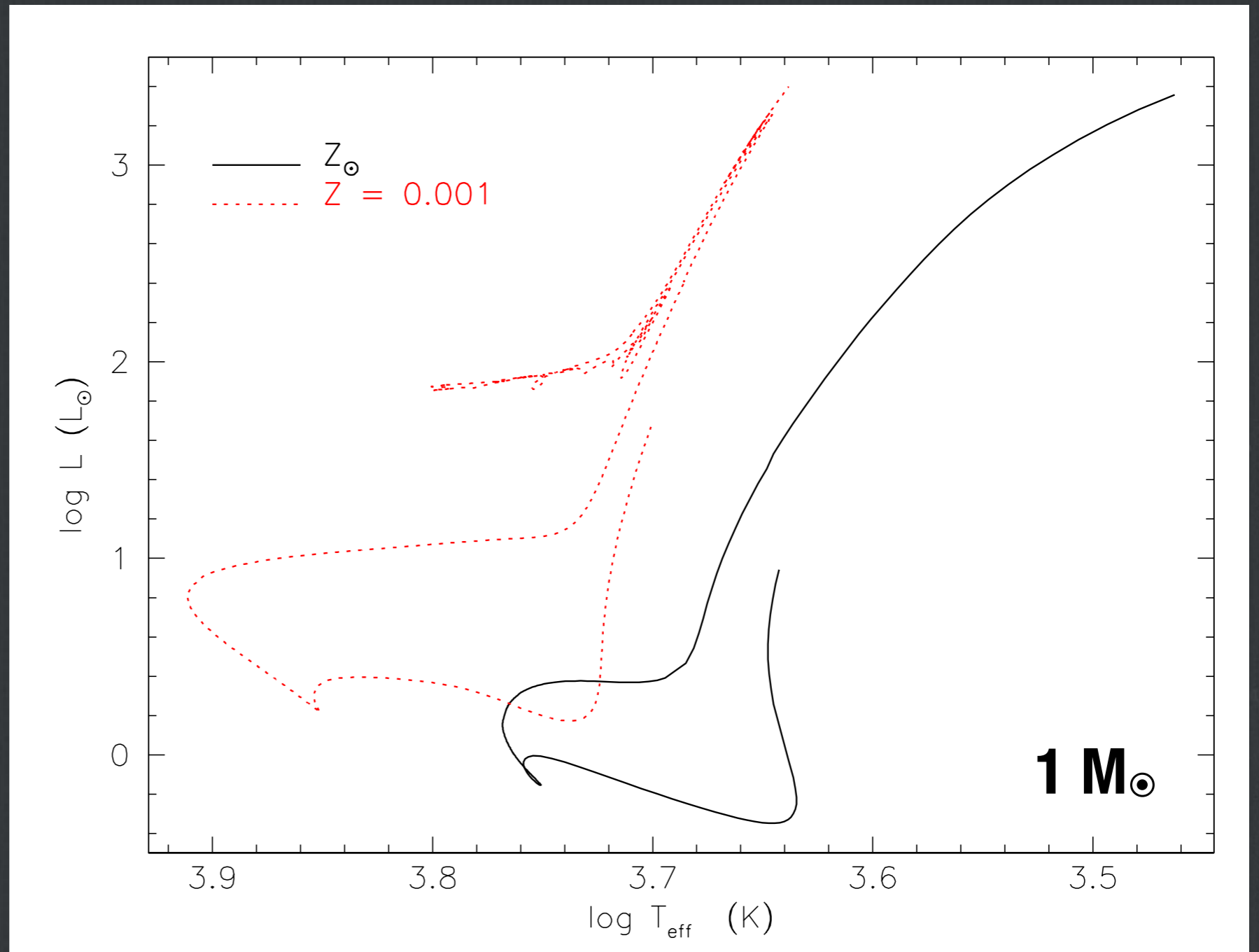
# Habitable zone

## Metallicity

shift in  $T_{\text{eff}}$  and  $L_*$   
due to opacity  
effects

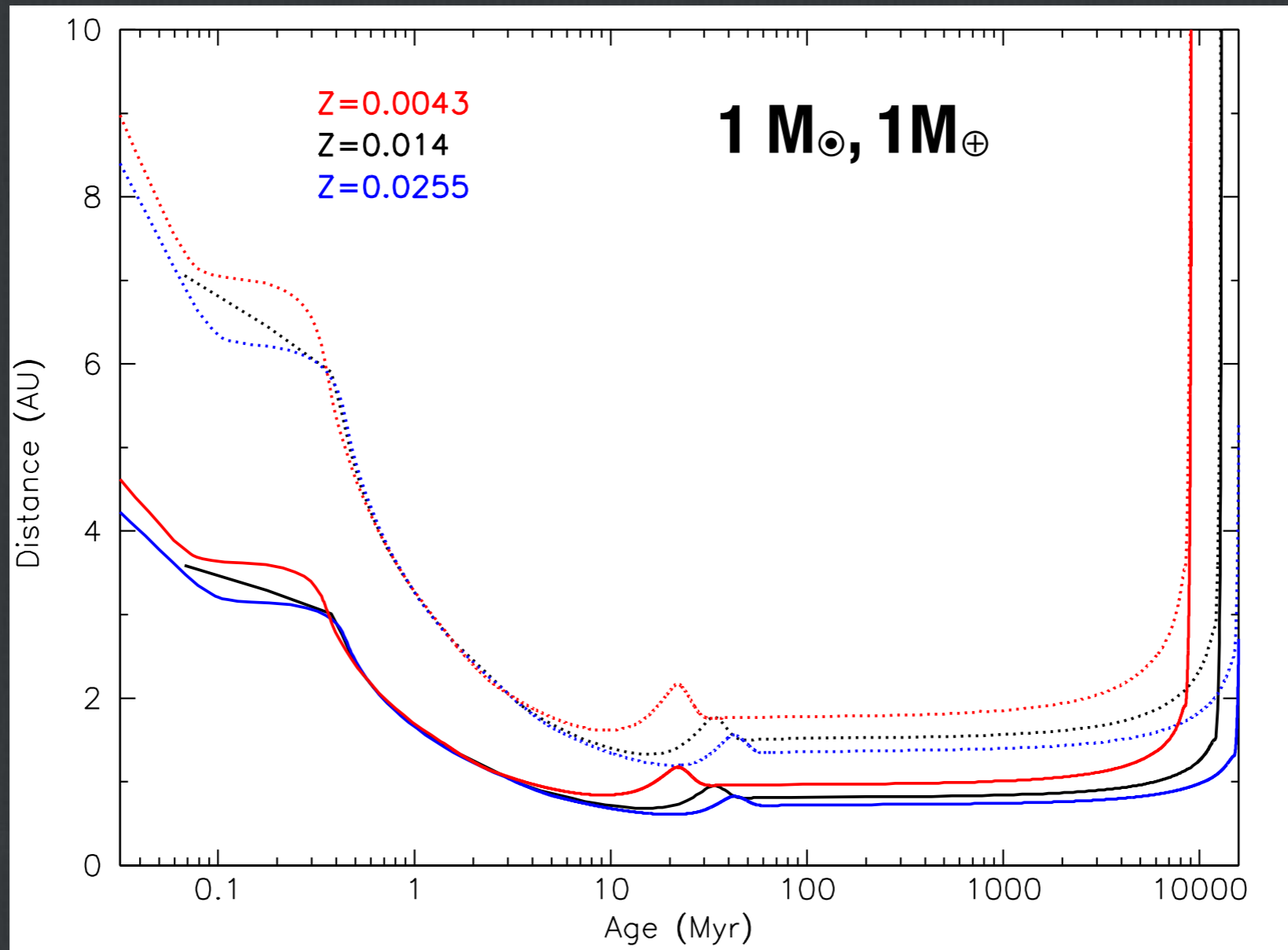


strong impact on HZ



# Habitable zone

## Metallicity



- **Increases** towards lower metallicities
- **Correct** estimation of metallicity = **crucial**

# Habitable zone

## Impact of rotation

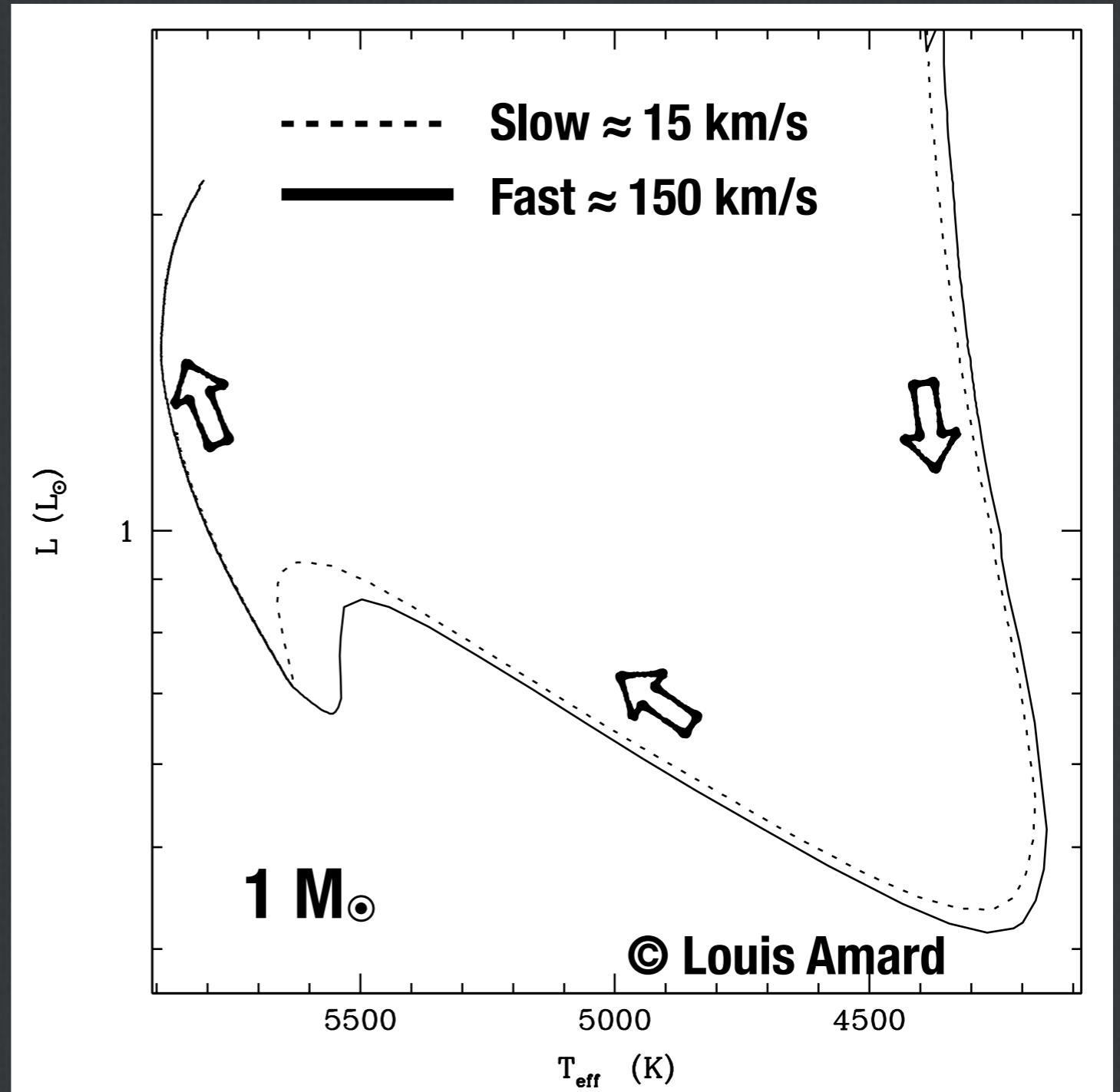
### Centrifugal effect

$$R_* \nearrow \rightarrow T_{\text{eff}} \searrow$$

shift of 150 K in  $T_{\text{eff}}$   
but  
 $L_*$  almost the same

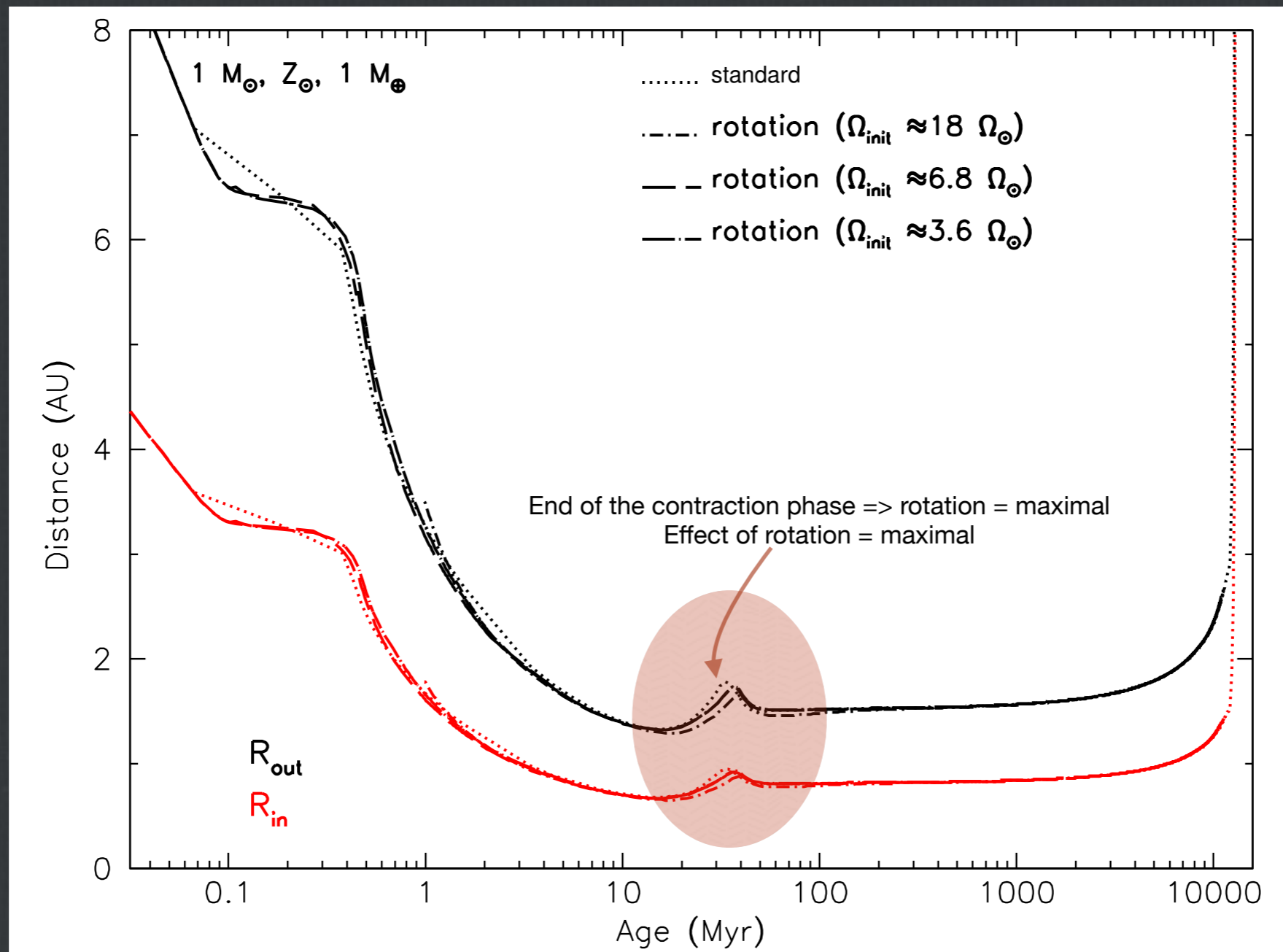


small impact on HZ



# Habitable zone

## Impact of rotation



Rotating model from Amard et al. (in prep.)

# Habitable zone

## Stellar activity

$Z_{\odot}$      $1 M_{\odot}$     Rotation

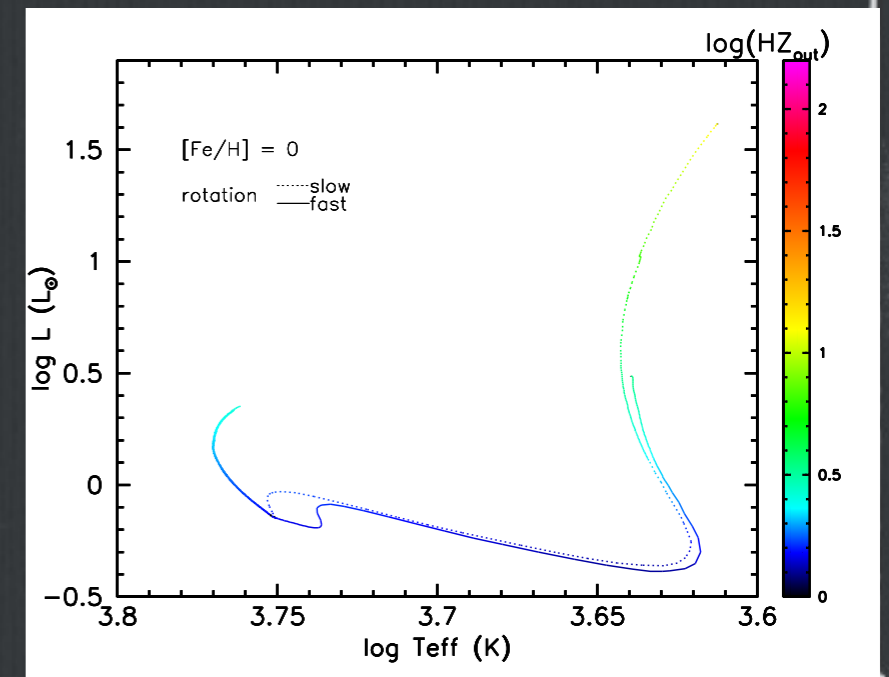
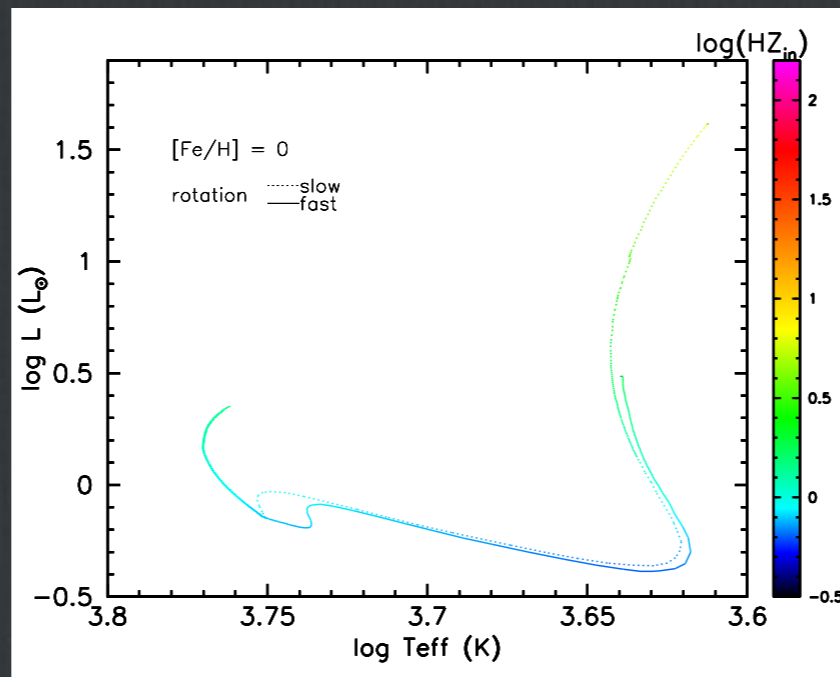
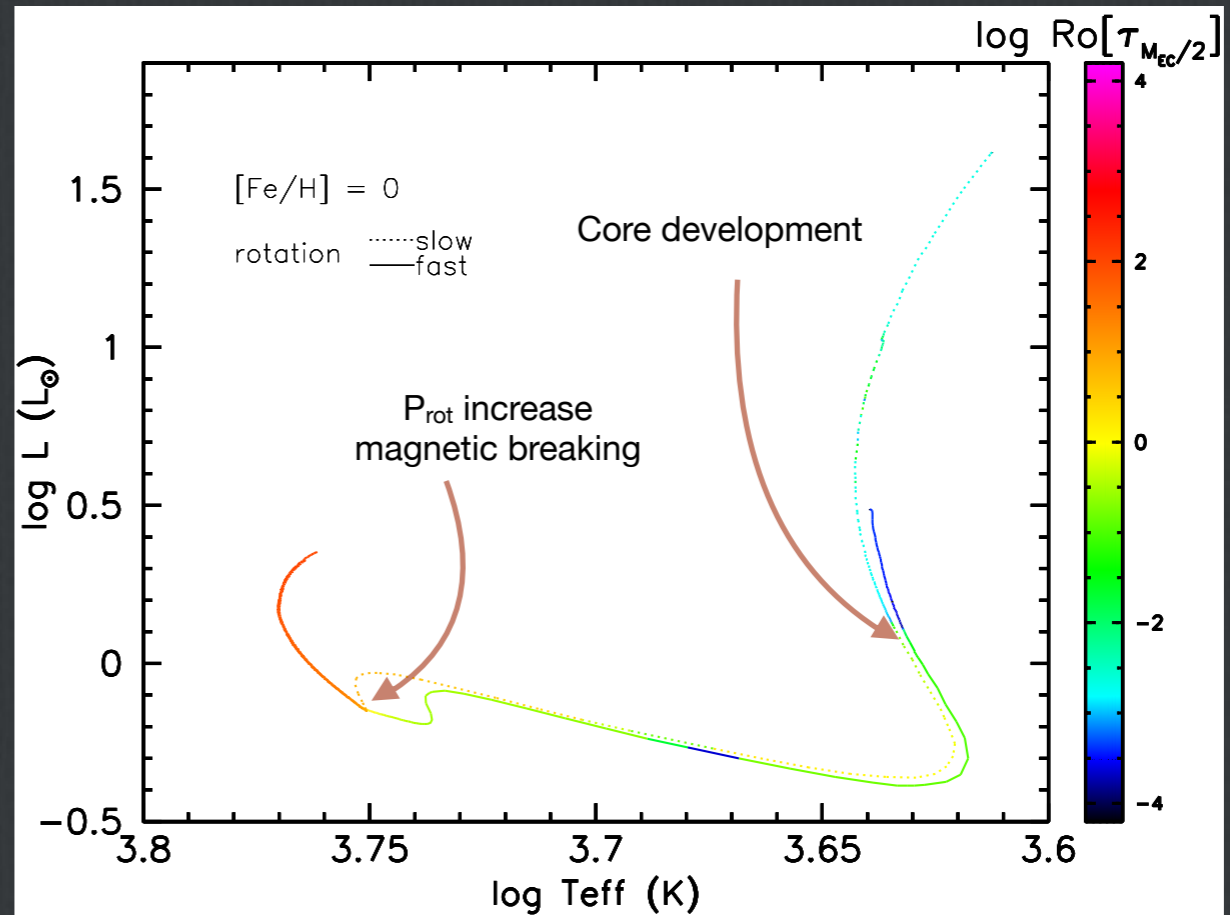
$$Ro = \frac{P_{rot}}{\tau_{conv}}$$

$Ro < 1 = \text{stellar activity}$

High stellar activity  
during the PMS  
when HZL **closest** for  
from the star



Impact on emergence/  
complexification of life?



# Habitable zone

CHZ

Gallet et al. (in prep.)

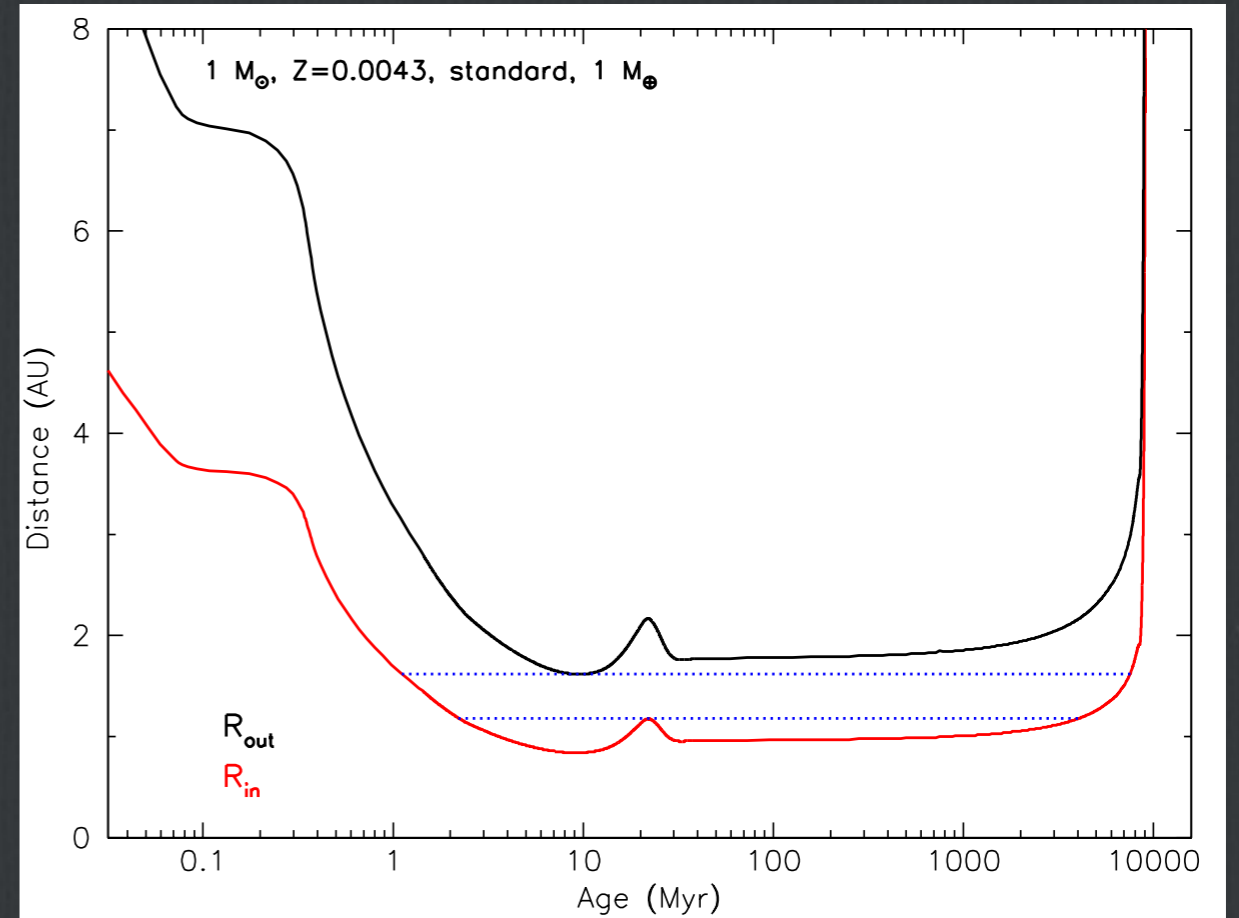
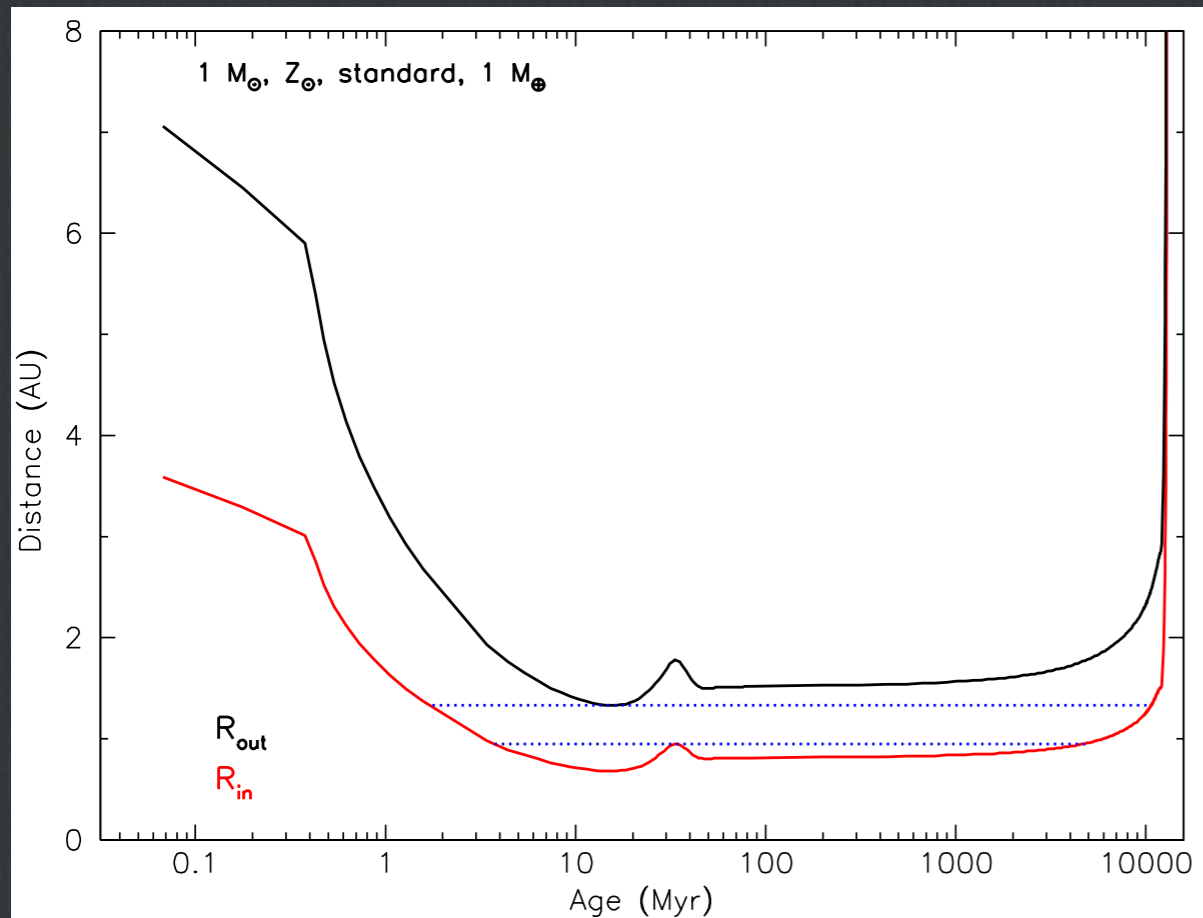
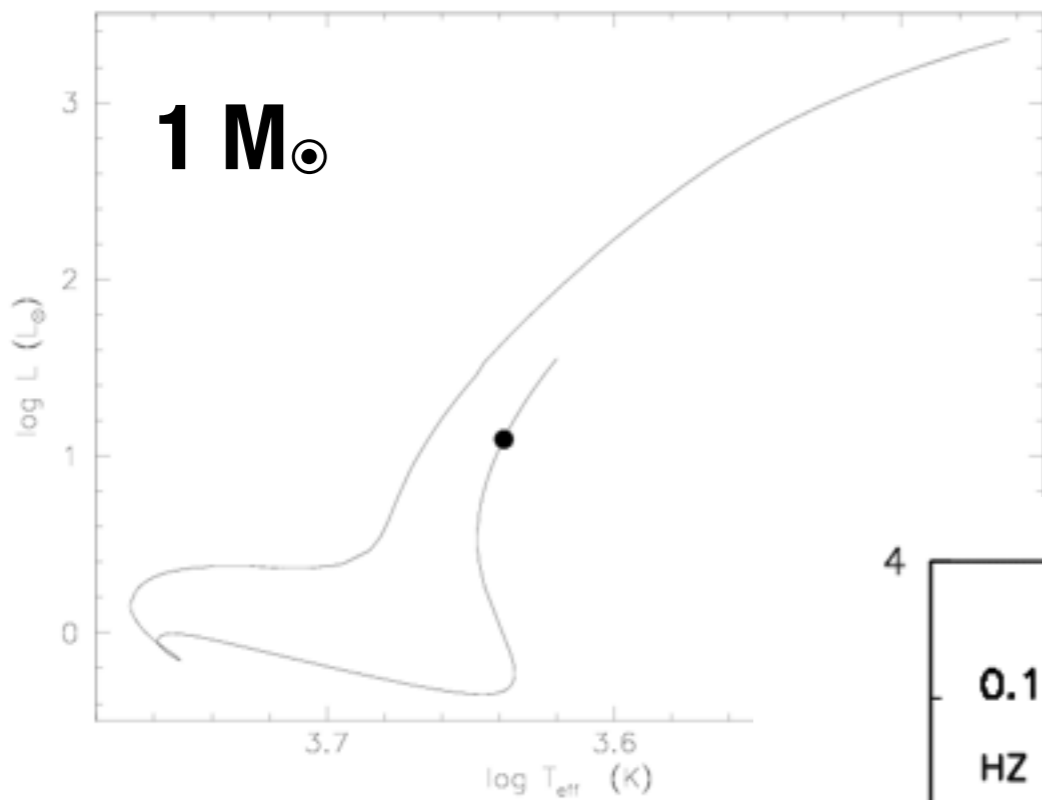


Table 2. Size of the CHZ as a function of stellar mass for solar metallicity.

Mass	HZ <sub>in</sub> (AU)	Start (Myr)	Time (Gyr)	HZ <sub>out</sub> (AU)	Start (Myr)	Time (Gyr)
0.5 M <sub>⊙</sub>	0.22	79.5	19.35	0.41	10.80	19.41
1 M <sub>⊙</sub>	0.95	3.74	4.77	1.33	1.76	10.67
1.5 M <sub>⊙</sub>	2.34	0.917	1.87	2.53	0.721	2.19

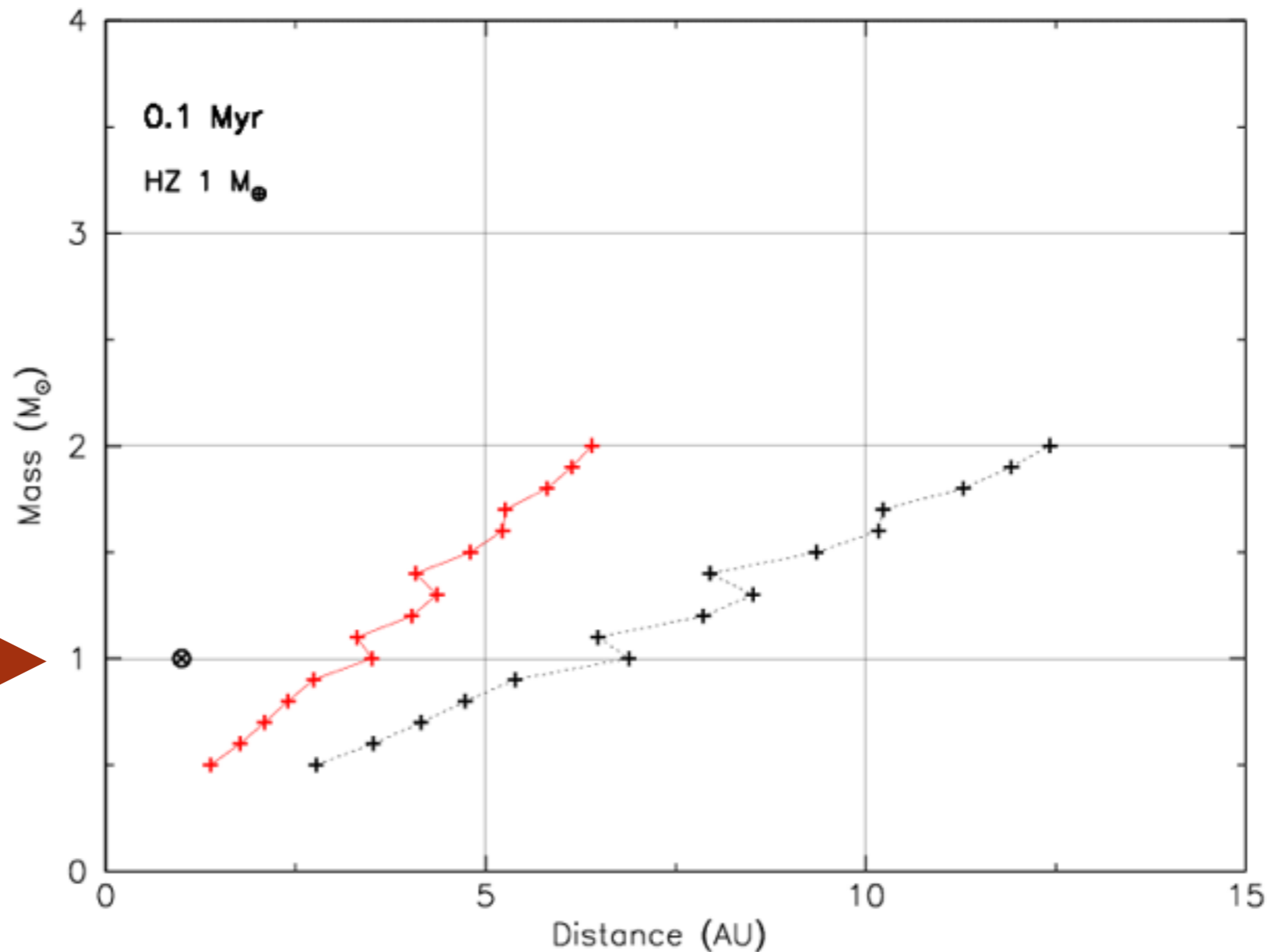
Start (Myr) = starting age of the CHZ, Time (Gyr) = duration of the CHZ

- **Longer** for planet close to the outer edge
- **Decrease** towards lower metallicity
- **Longer** for lower mass stars



Z<sub>⊙</sub>      0.5-2 M<sub>⊙</sub>      No rotation

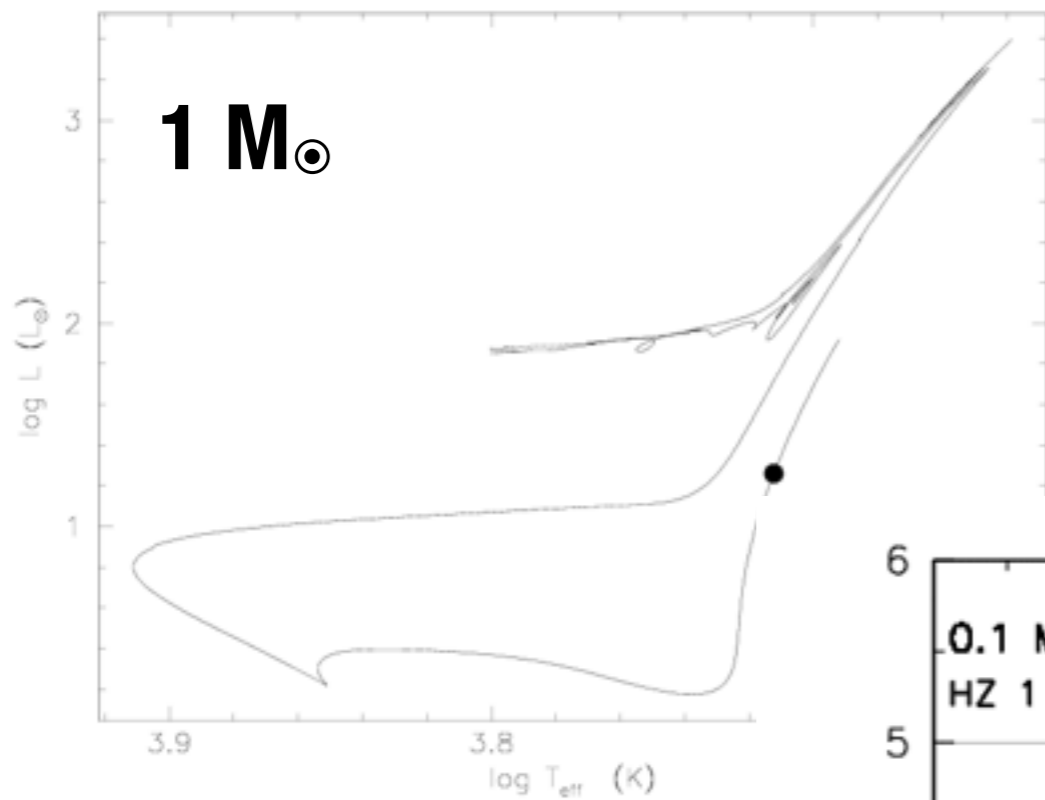
Kopparapu et al. (2014) + STAREVOL



Sun's age  
**0.95 - 1.76 AU**

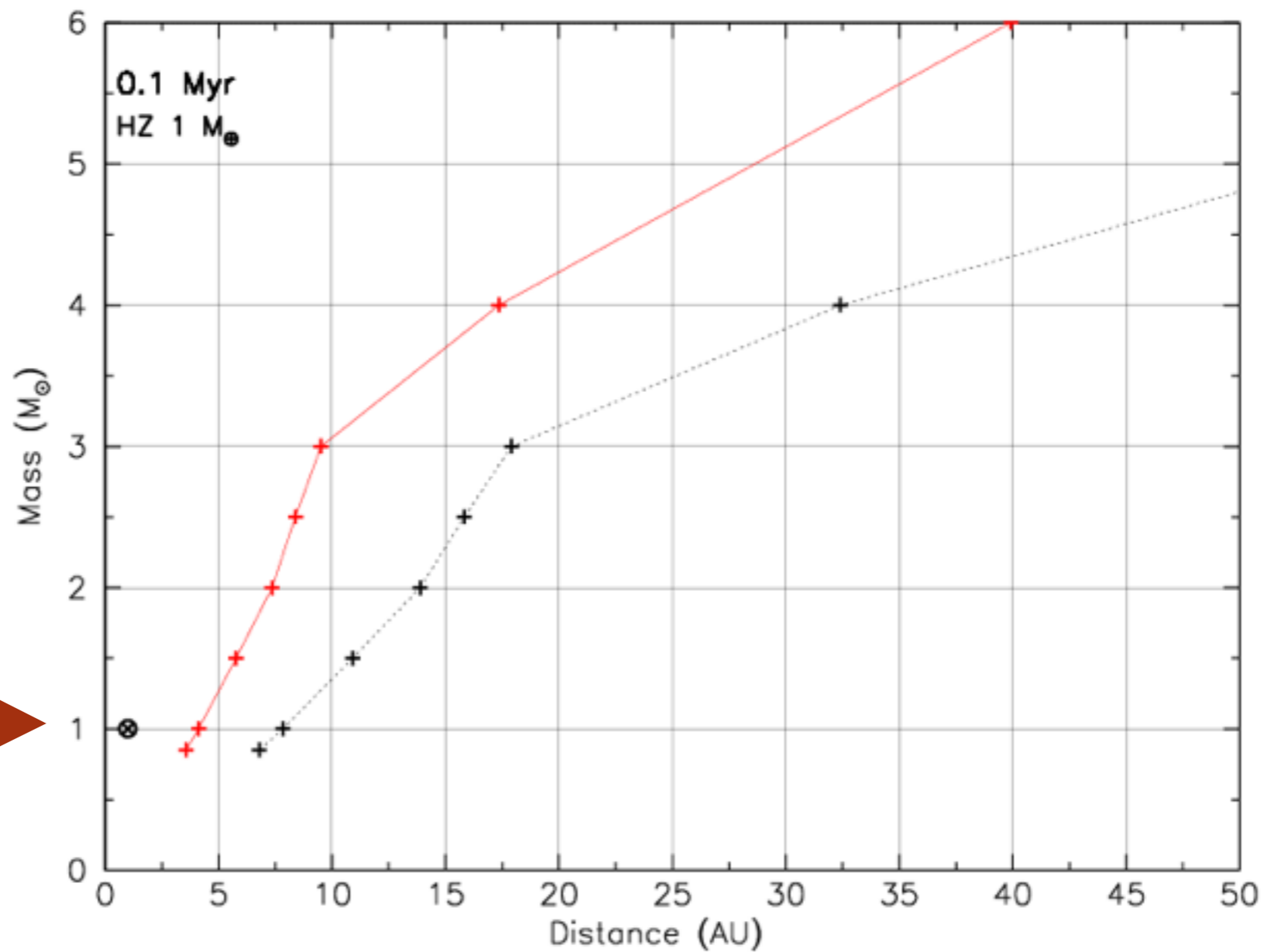
Earth Habitability  
**3 Myr - 6 Gyr**

**1 M<sub>⊙</sub>**



**Z=0.001    0.8-6 M<sub>⊙</sub>    No rotation**

**Kopparapu et al. (2014) + STAREVOL**





# Conclusion/Perspective

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- Mass and metallicity control HZ evolution**
    - => require precise estimation of  $M_*$  and Fe/H
  - Stellar models should be used to get HZ evolution**
    - estimation of CHZ / duration of planet inside HZ?
- 
- Provide the community with an online tool to estimate habitability**
  - Couple atmospheric model to STAREVOL to directly get the flux receive by planet**
  - Include tidal interaction/dissipation in STAREVOL => Stéphane Mathis @ CEA Saclay**
  - Magnetic interaction? => Aline Vidotto @ Geneva**

[obswww.unige.ch/Recherche/evol/starevol/HZcalculator.php](http://obswww.unige.ch/Recherche/evol/starevol/HZcalculator.php)

STELLAR EVOLUTION

Home > Research > Habitability > HZ calculator

Members  
Research  
Database  
Publications

References

Amard et al. (in prep.)  
Kopparapu et al. (2013)  
Kopparapu et al. (2014)  
Selsis et al. (2007)  
Underwood et al. (2003)

Calculation of the Habitable Zones

Habitable Zone

Days of year relative to Earth

Radius of orbit relative to Earth

Venus  
Earth  
Mars

Stellar models from Amard et al. (in prep.)

Please enter the stellar mass and chose your favorite HZ prescription

Rotation ? (not implemented yet)  
Yes   
No

No rotation  
 Kopparapu  Selsis  Underwood

Stellar mass = 1.0

Planet location (AU) ? 2

Age (yr) ? 1e6

CHZ ?  yes  no

Selected mass = 1  $M_{\odot}$   
Metallicity =  $Z_{\odot}$   
Prescription = Kopparapu et al. 2014

Red cross = outside the habitable zone  
Blue cross = inside the habitable zone

Time in HZ<sub>in</sub> 3.74 Myr - Time out HZ<sub>in</sub> 4.77 Gyr  
Time in HZ<sub>out</sub> 1.76 Myr - Time out HZ<sub>out</sub> 10.67 Gyr  
HZ<sub>in</sub> 0.95 AU - HZ<sub>out</sub> 1.33 AU

Online tool already available

we will provide stellar grids  
and other  
visualisation tools



stay tuned