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**UNIVERSITÄT
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EEG
**EXOPLANETS
& EXOCLIMES
GROUP**

The development of an open-source chemical kinetics: **VULCAN**

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- open-source Python code
- reduced network
- flexibility

EQ v.s. Kinetics

Equilibrium chemistry

Minimise Gibbs free energy
(only final products relevant)

Deals with net reactions only

Requires tabulated database
of Gibbs free energies of
formation

Does not consider atmospheric
circulation or photochemistry

Chemical kinetics


Solve set of mass conservation
equations (pathways relevant)

Deals with transient reactions
and radicals

Requires tabulated rate coefficients
for both forward and reverse
reactions

May mimic atmospheric circulation
using eddy diffusion

Hierarchical Chemical Network



	Equilibrium thermochemistry	Disequilibrium: vertical mixing	future:
H	started in 2015		Disequilibrium: photochemistry & condensation
H + O			
H + O + C		● where we are	
H + O + C + N		by the end of 2015	

chemical complexity

Model Description

- ❖ Reduced network (25 species & ~300 reactions)
(H,H₂,O,OH,H₂O,CH,C,CH₂,CH₃,CH₄,C₂,C₂H₂,C₂H₃,C₂H,C₂H₄,C₂H₅,
C₂H₆,C₄H₂,CO,CO₂,CH₂OH,H₂CO,HCO,CH₃O,CH₃OH)
- ❖ Group mechanism
- ❖ Rate coefficient from NIST / reversed
- ❖ solving stiff ODEs: $\frac{dY}{dt} = f(Y, t) = P - L$
- ❖ semi-implicit Euler solver (analytical Jacobian matrix)
$$f(Y^{n+1}, t^{n+1}) \approx f(Y^n, t^n) + \left. \frac{\partial f}{\partial Y} \right|_{Y^n} \cdot (Y^{n+1} - Y^n)$$
- ❖ Crank-Nicolson method for diffusion
- ❖ Challenge: stiffness & coding time

Auto-generated script

```
# C/H/O chemistry

#R1
H + H + M -> H2 + M
#R3
H + O + M -> OH + M
#R5
H + H2O -> OH + H2
#R7
O + H2 -> OH + H
#R9
O + H2O -> 2*OH
#R11
OH + H + M -> H2O + M
#R13
H + CH -> H2 + C
#R15
H + CH + M -> CH2 + M
#R17
H + CH2 -> CH + H2
#R19
CH2 + H2 -> H + CH3
#R21
H + CH3 + M -> CH4 + M
#R23
H + CH4 -> CH3 + H2
#R25
C + CH -> C2 + H
#R27
CH3 + C -> H + C2H2
```

❖ flexibility
of the
model

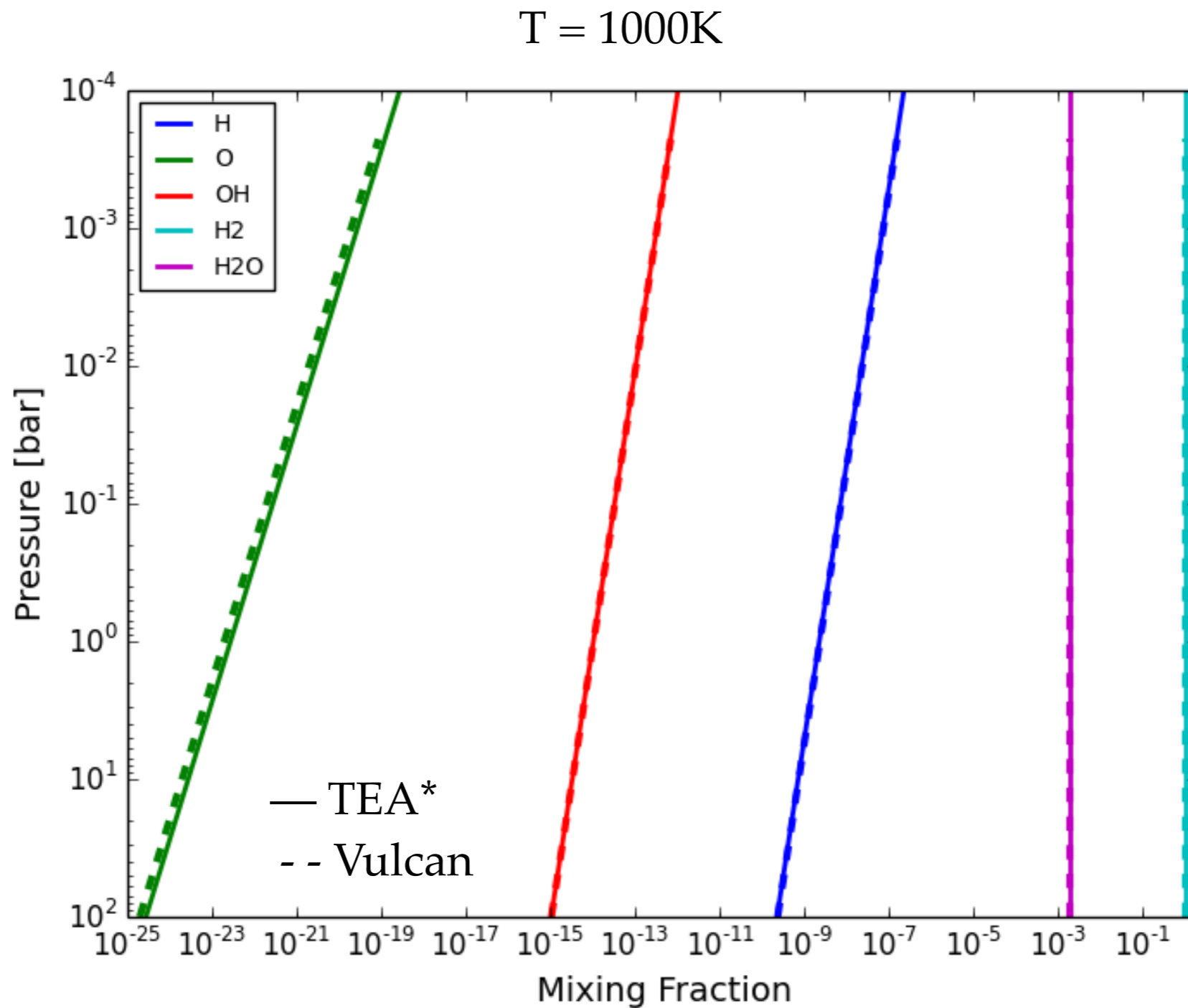
inspired from Kinpy

```
## Mapping ##
H: y[0], H2: y[1], O: y[2], OH: y[3], H2O: y[4], CH: y[5], C: y[6],

H 0 -1*v_1(y[1], y[0]) -1*v_1(y[1], y[0]) -1*v_3(y[0], y[2], y
H2 1 +1*v_1(y[1], y[0]) +1*v_5(y[1], y[0], y[4], y[3]) -1*v_7(y
O 2 -1*v_3(y[0], y[2], y[3]) -1*v_7(y[1], y[0], y[2], y[3]) -1
OH 3 +1*v_3(y[0], y[2], y[3]) +1*v_5(y[1], y[0], y[4], y[3]) +1
H2O 4 -1*v_5(y[1], y[0], y[4], y[3]) -1*v_9(y[4], y[2], y[3]) +1
CH 5 -1*v_13(y[1], y[0], y[6], y[5]) -1*v_15(y[0], y[7], y[5])
C 6 +1*v_13(y[1], y[0], y[6], y[5]) -1*v_25(y[10], y[6], y[5])
CH2 7 +1*v_15(y[0], y[7], y[5]) -1*v_17(y[1], y[0], y[7], y[5])
CH3 8 +1*v_19(y[1], y[0], y[7], y[8]) -1*v_21(y[0], y[8], y[9])
CH4 9 +1*v_21(y[0], y[8], y[9]) -1*v_23(y[1], y[0], y[8], y[9])
C2 10 +1*v_25(y[10], y[6], y[5], y[0]) +1*v_65(y[10], y[6]) -1*v
C2H2 11 +1*v_27(y[0], y[6], y[8], y[11]) -1*v_29(y[0], y[12],
C2H3 12 +1*v_29(y[0], y[12], y[11]) -1*v_35(y[1], y[0], y[12],
C2H 13 -1*v_31(y[1], y[0], y[13], y[11]) -1*v_55(y[0], y[13], y[1
C2H4 14 +1*v_37(y[0], y[14], y[12]) +1*v_39(y[1], y[0], y[14],
C2H5 15 +1*v_45(y[0], y[15], y[14]) -1*v_47(y[0], y[15], y[8])
C2H6 16 +1*v_51(y[0], y[15], y[16]) -1*v_53(y[1], y[0], y[15],
C4H2 17 +1*v_55(y[0], y[13], y[17], y[11])
CO 18 -1*v_57(y[0], y[19], y[18], y[3]) +1*v_63(y[1], y[4], y[19
CO2 19 +1*v_57(y[0], y[19], y[18], y[3]) -1*v_63(y[1], y[4], y[19
CH2OH 20 +1*v_167(y[20], y[8], y[0], y[3]) -1*v_169(y[20], y[21
H2CO 21 +1*v_169(y[20], y[21], y[0]) -1*v_171(y[1], y[0], y[21
HCO 22 +1*v_171(y[1], y[0], y[21], y[22]) -1*v_173(y[18], y[0], y
CH3O 23 +1*v_195(y[0], y[23], y[21]) -1*v_197(y[23], y[21], y[
CH3OH 24 +1*v_201(y[24], y[20], y[0]) +1*v_203(y[24], y[8], y[3
CH3CO 25 +1*v_223(y[25], y[11], y[3]) -1*v_267(y[8], y[25], y[1
...

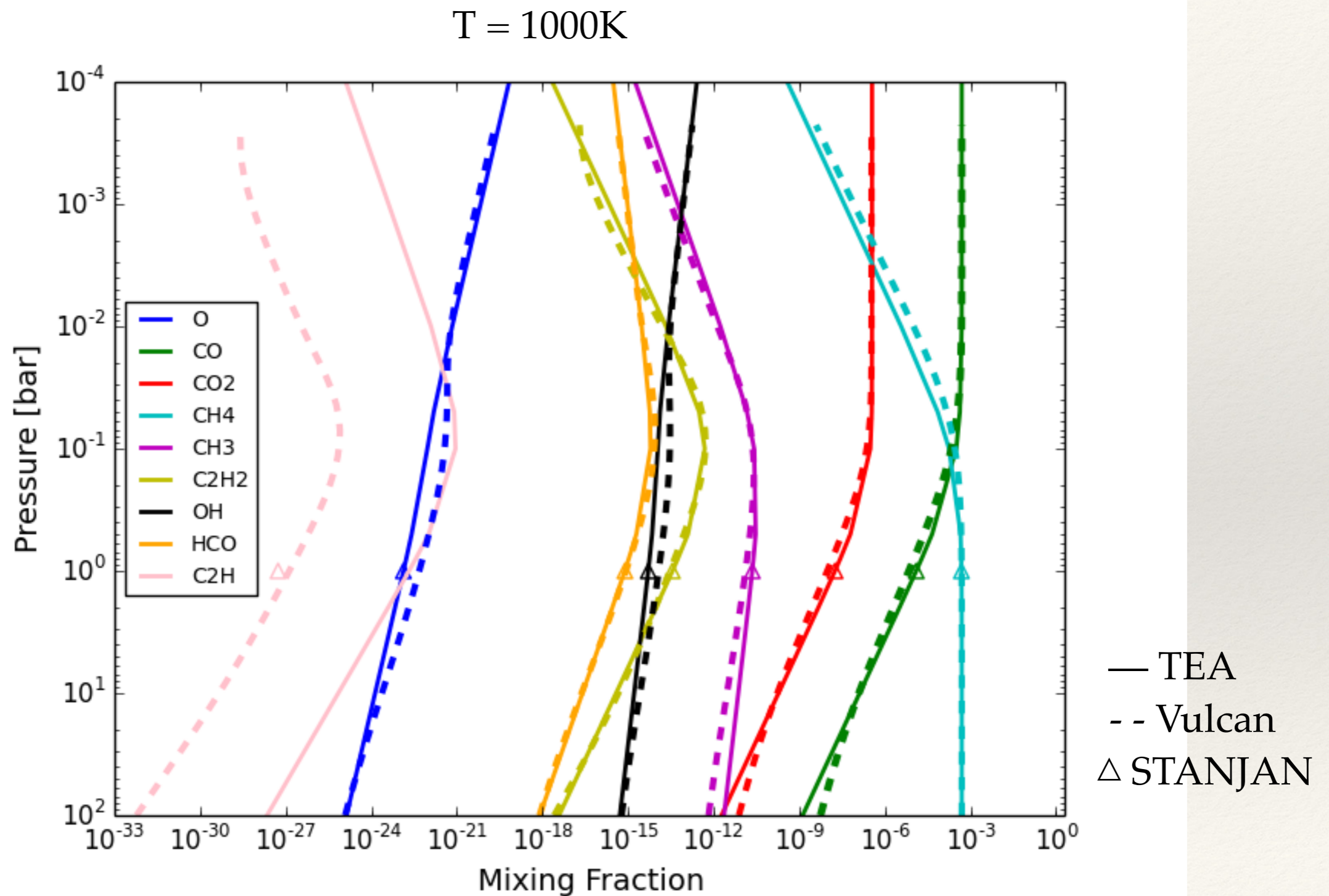
def chemdf(y):
    y = np.transpose(y)
    dydt = np.zeros(shape=y.shape)
    dydt[0] = -1*v_1(y[1], y[0]) -1*v_1(y[1], y[0]) -1*v_3(y[0], y
    dydt[1] = +1*v_1(y[1], y[0]) +1*v_5(y[1], y[0], y[4], y[3]) -1
    dydt[2] = -1*v_3(y[0], y[2], y[3]) -1*v_7(y[1], y[0], y[2], y[
    dydt[3] = +1*v_3(y[0], y[2], y[3]) +1*v_5(y[1], y[0], y[4], y[
    dydt[4] = -1*v_5(y[1], y[0], y[4], y[3]) -1*v_9(y[4], y[2], y[
```

Reproduce Chemical EQ



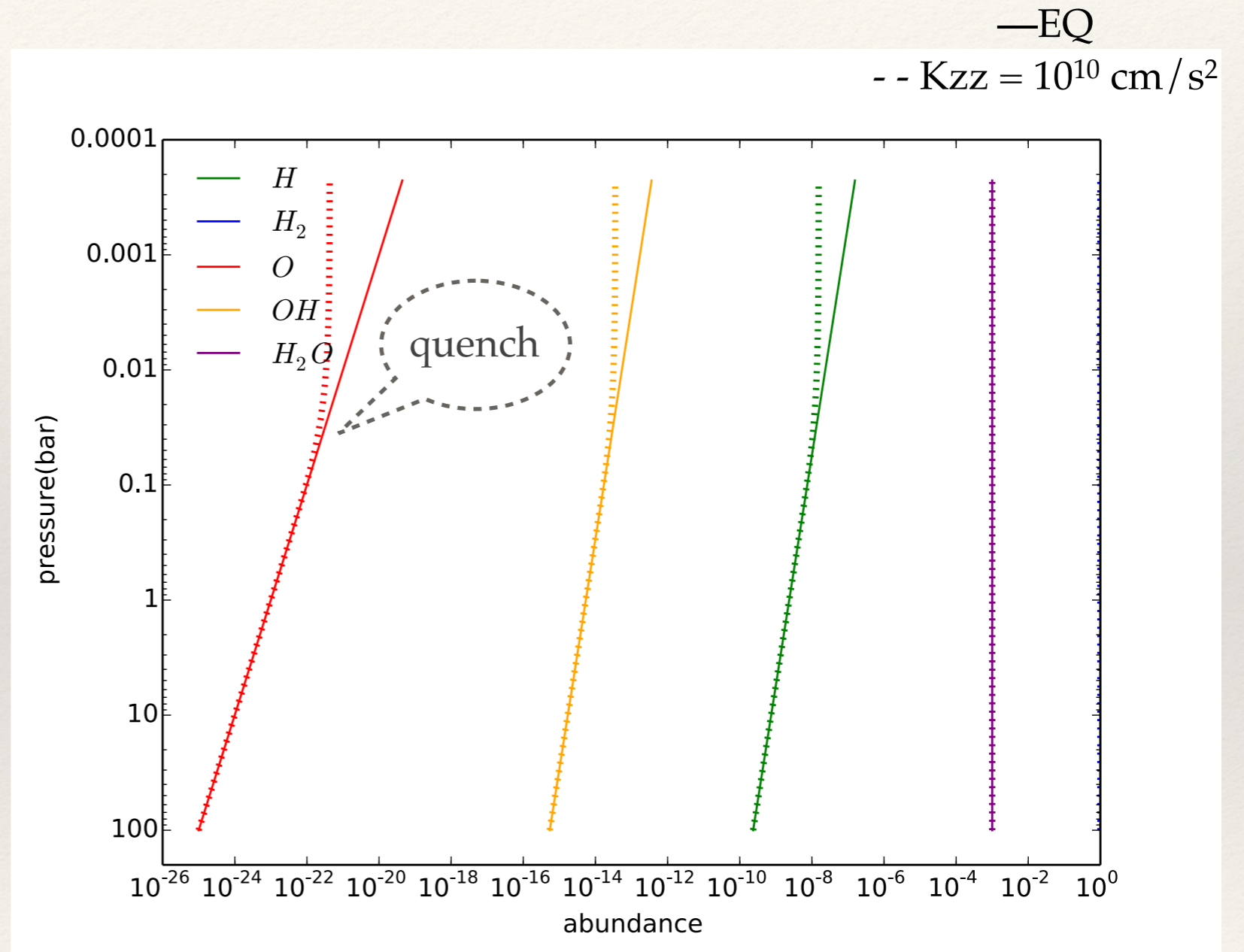
*Blecic, J., Harrington, J.,
& Bowman
arXiv:1505.06392v1

Reproduce Chemical EQ

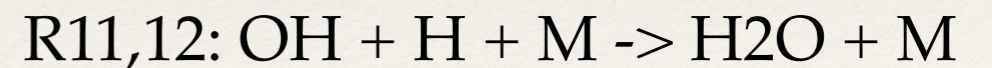
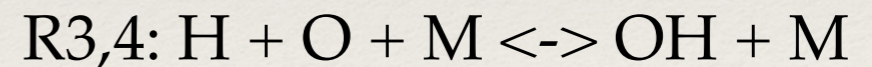
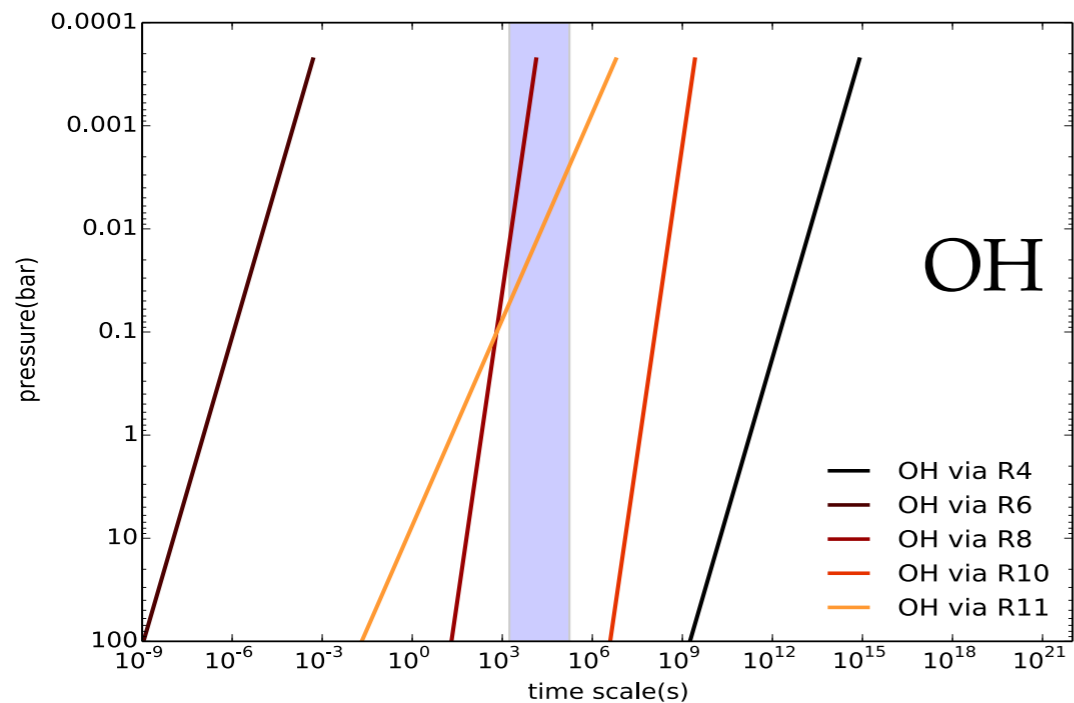
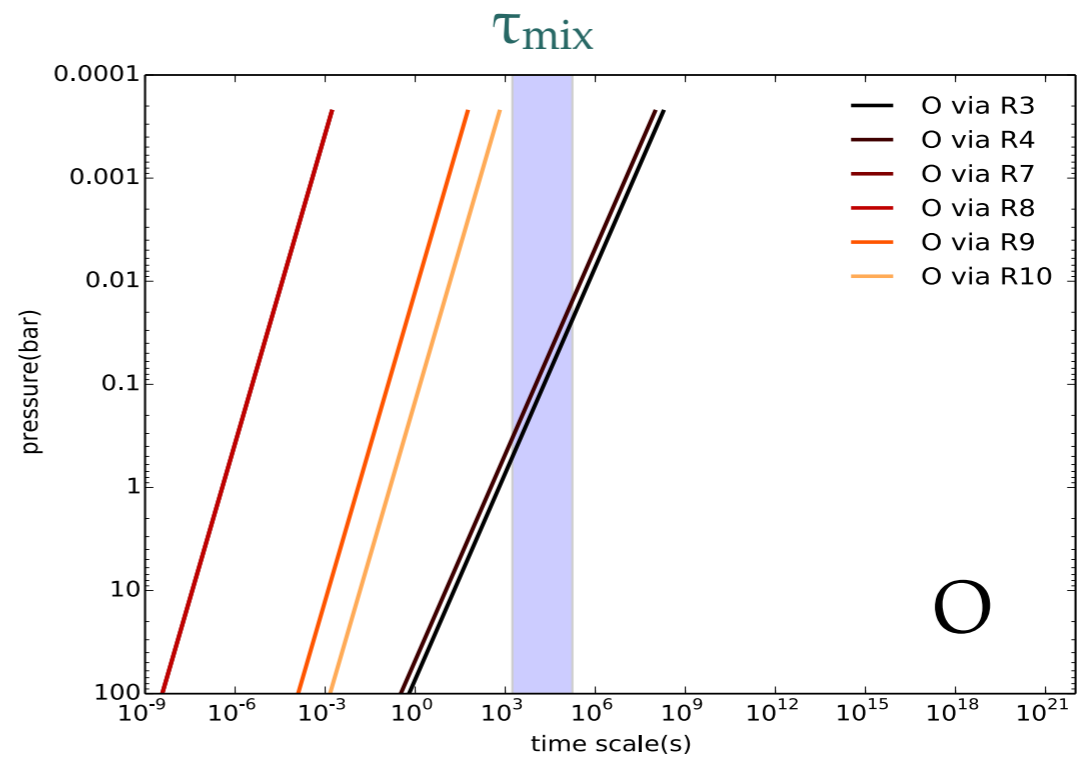
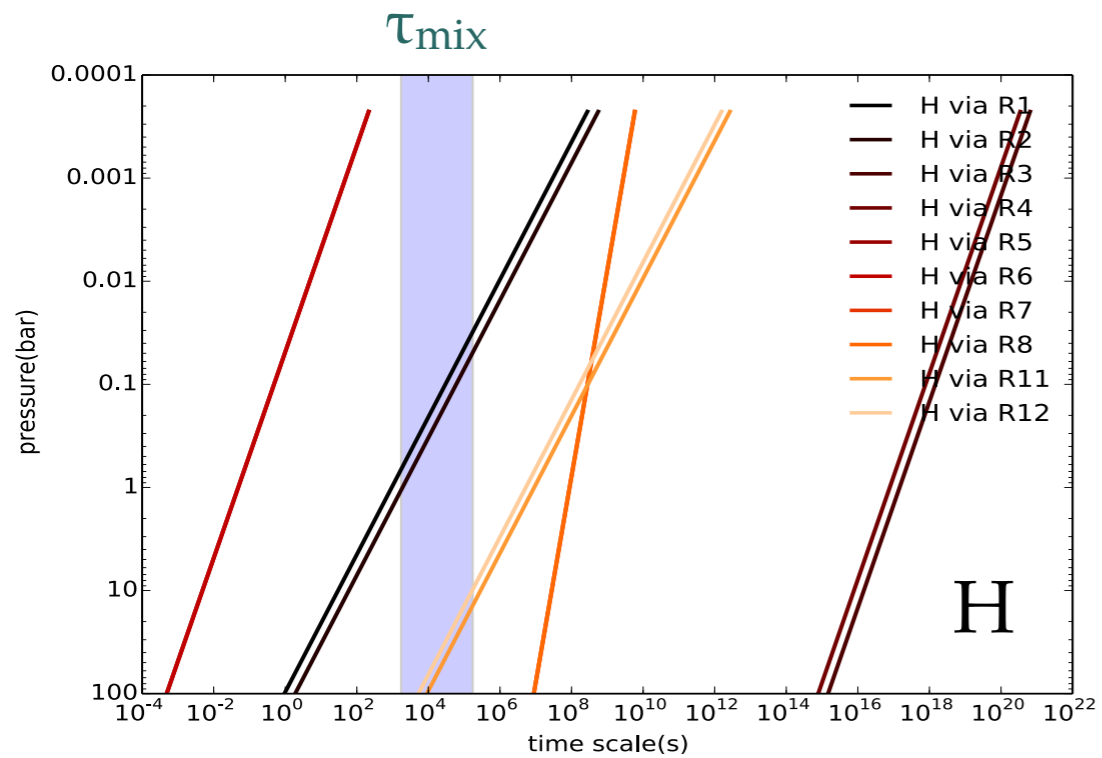


Transport-induced quench

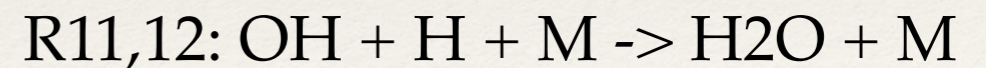
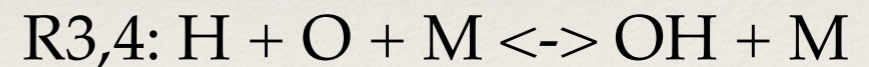
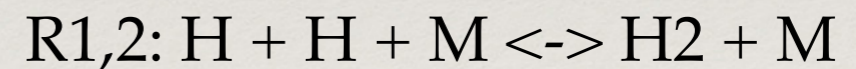
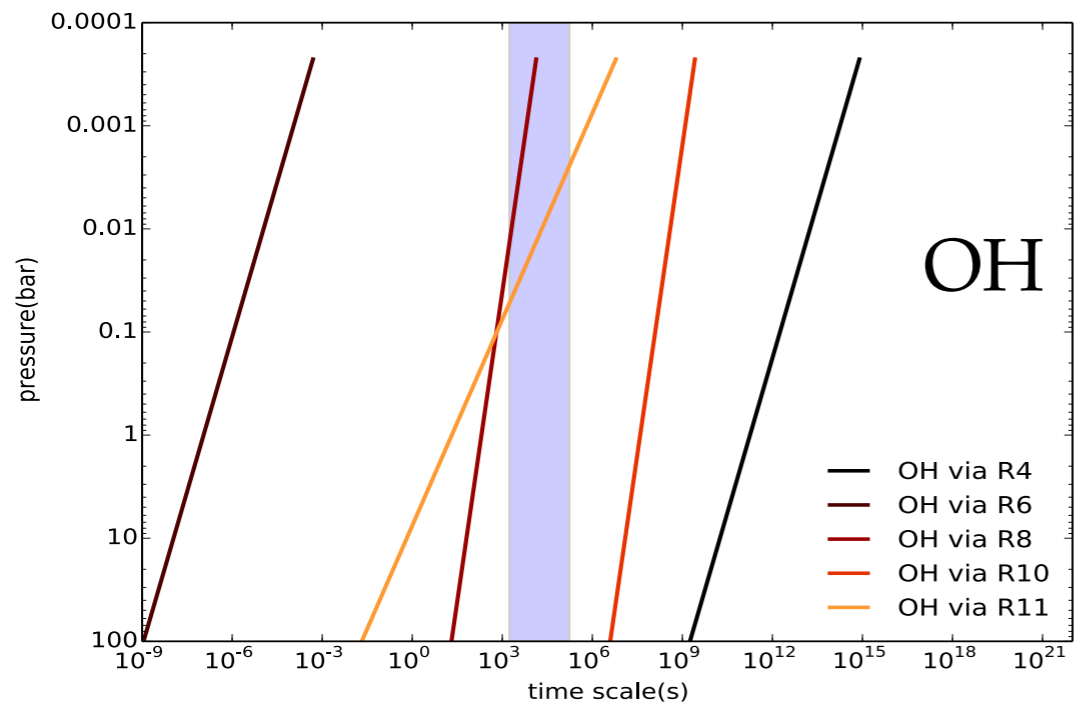
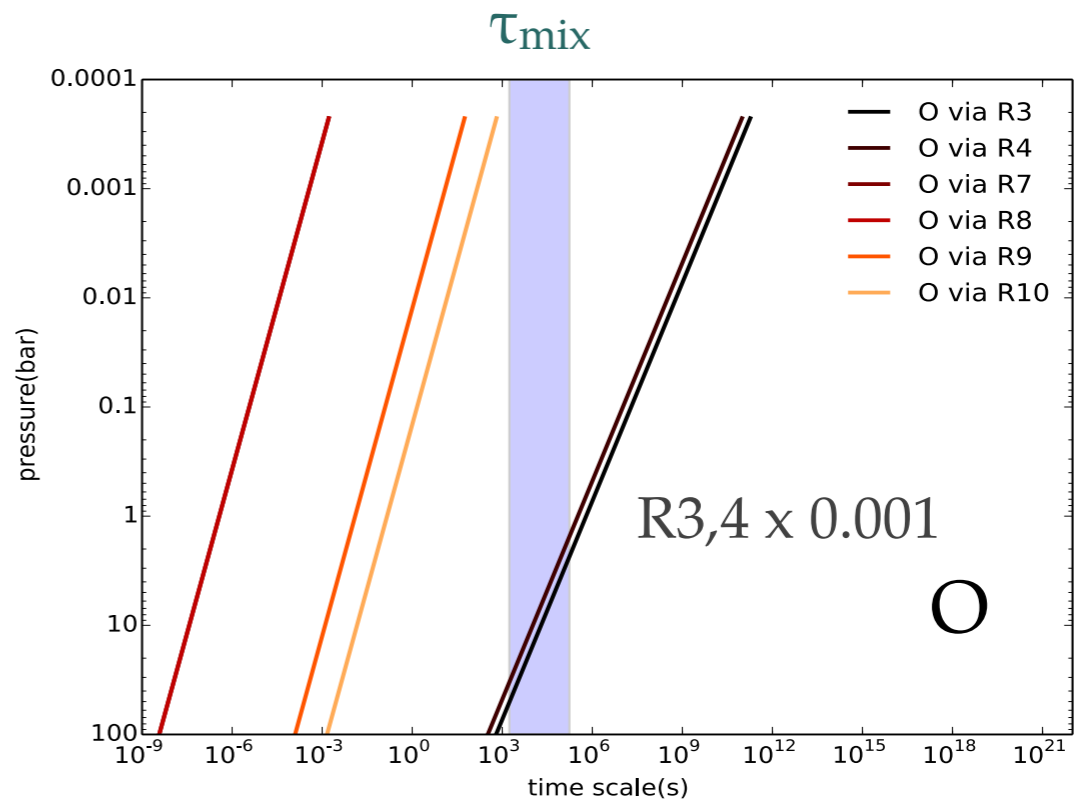
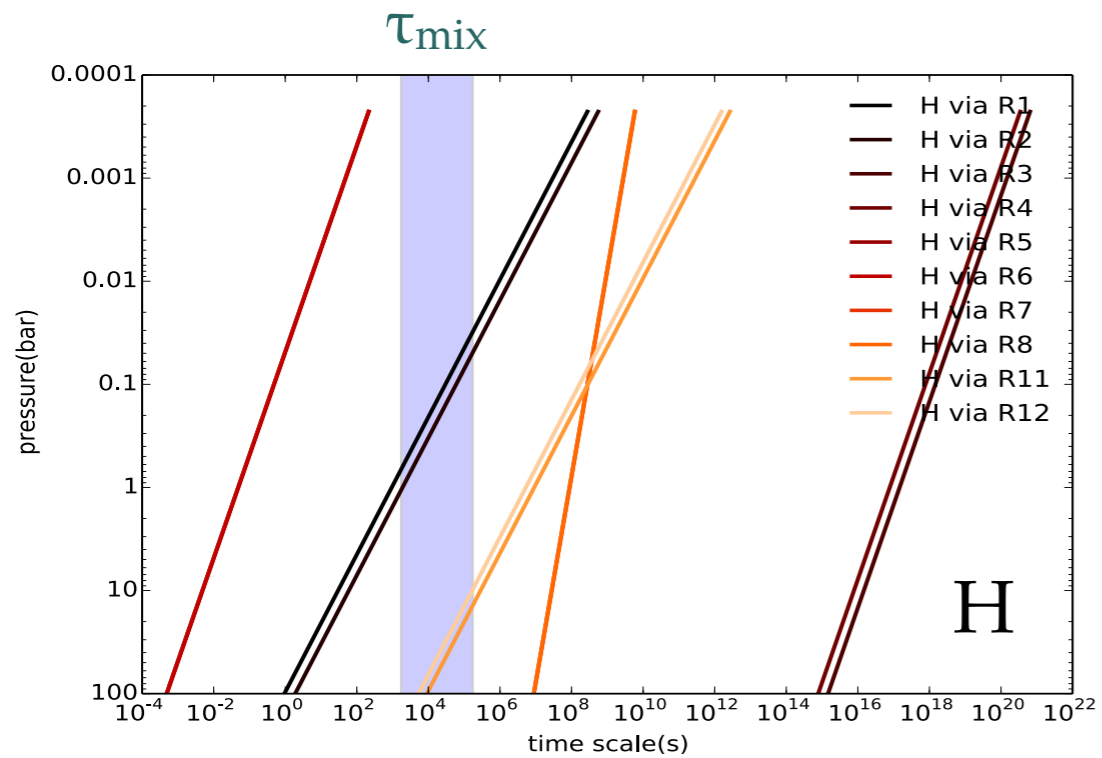
- ❖ $\tau_{\text{mix}} = L^2 / K_{\text{ZZ}}$
 $\tau_{\text{chem}} = [x] / (d[x] / dt)$
- ❖ quench level:
 $\tau_{\text{chem}} = \tau_{\text{mix}}$
- ❖ Identifying the rate-limiting step is crucial!



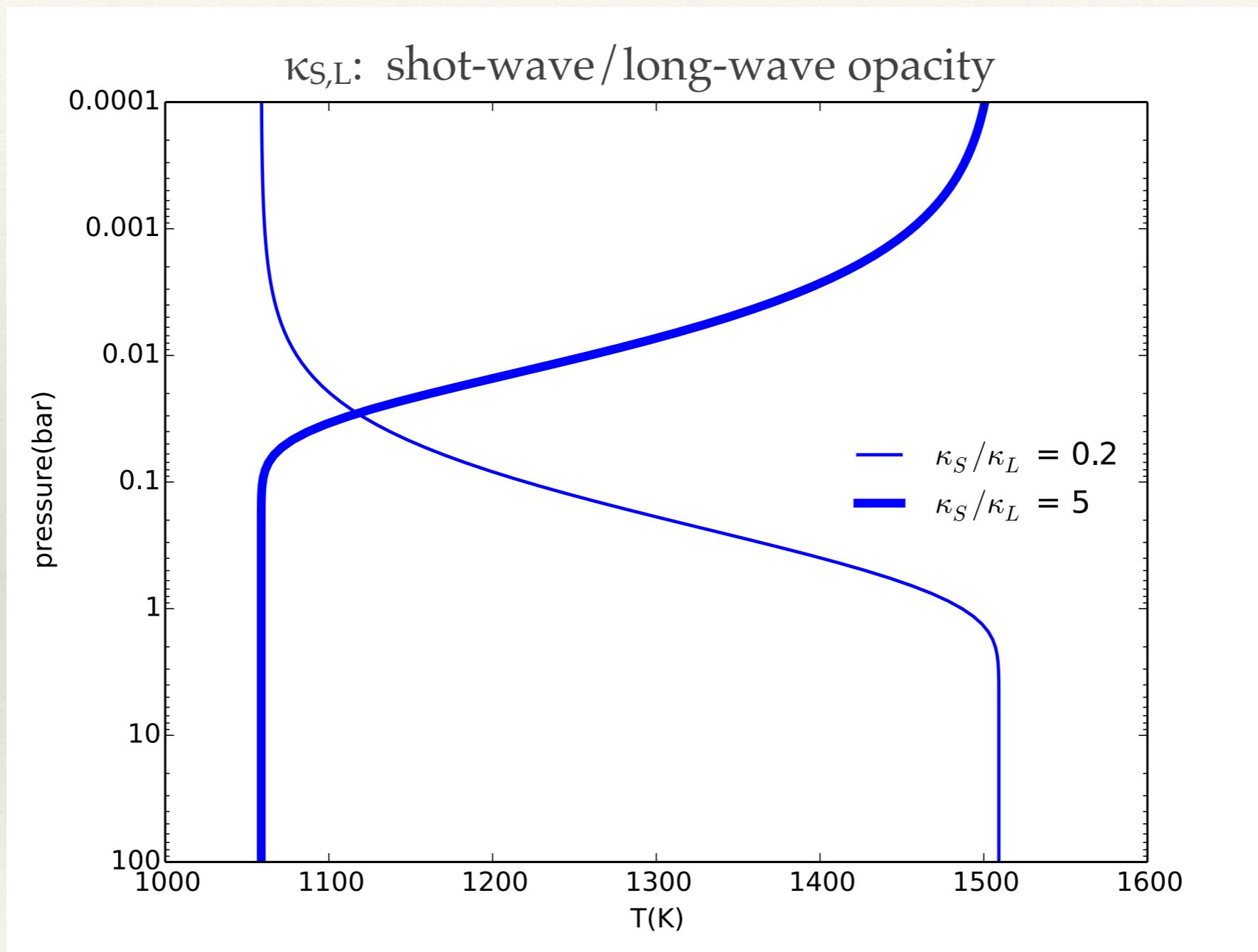
Transport-induced quench



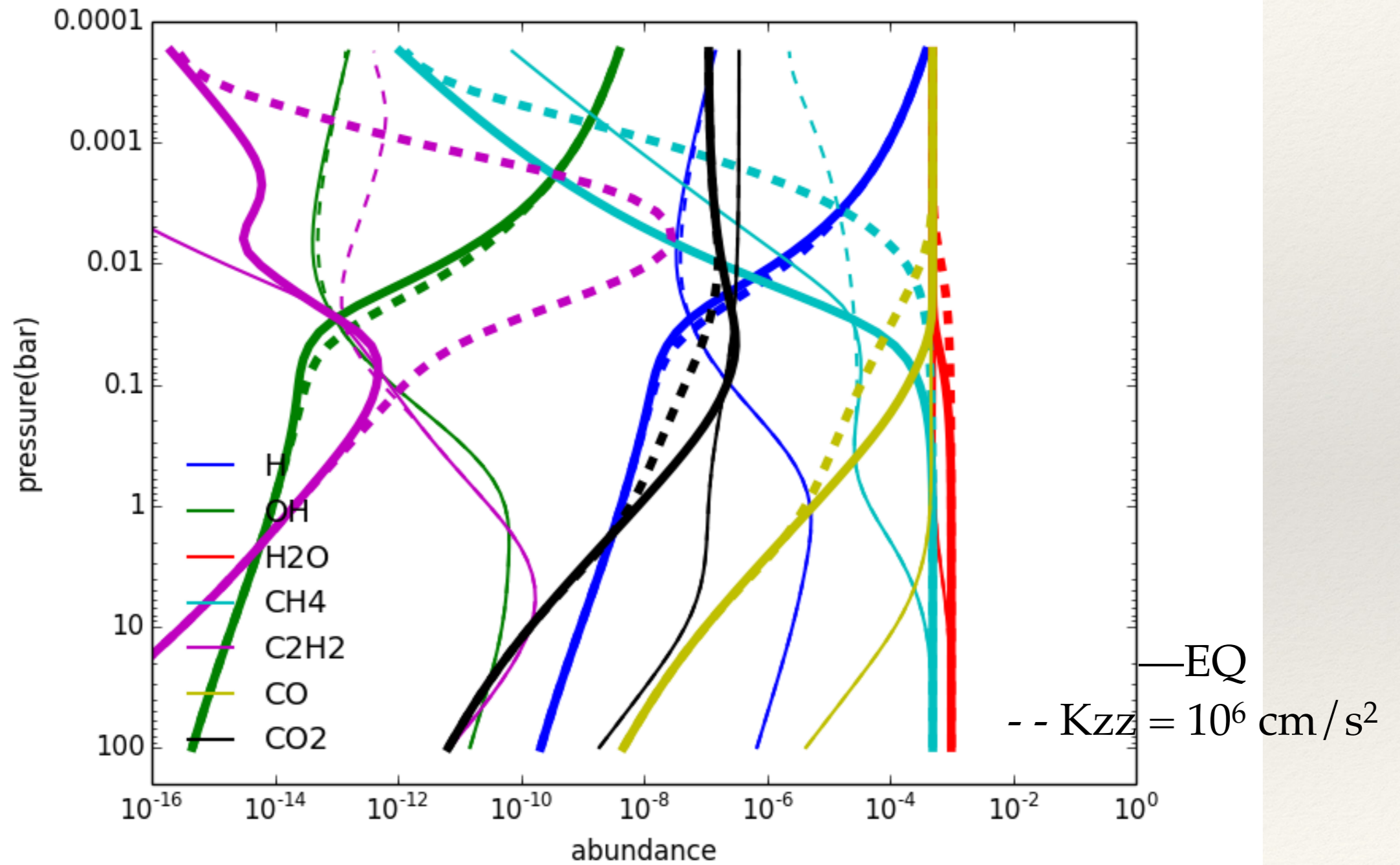
Transport-induced quench



Test on TP-profile



Test on TP-profile





ESP

EXOCLIMES SIMULATION PLATFORM

Our dream: an open-source set of modern computer codes, established as community tools

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Shang-Min Tsai, Maria Oreshenko,
Joao Mendonca, Simon Grimm,
Daniel Kitzmann, Frank Wagner
Collaborators: Sid Mishra, Roger Kaeppli,
Jim Lyons, Adam Burrows, Mike Line

Professor Kevin Heng
University of Bern, Switzerland

- The Exoclimes Simulation Platform (ESP) is a dream and vision to provide a versatile set of codes, for understanding exoplanetary atmospheres, to the scientific community.
- The ESP has three main components: HELIOS (radiation), VULCAN (chemistry) and THOR (fluid dynamics).