LIFE BASED ON METHANE ON SATURN'S MOON TITAN ?

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Introduction: what do we knew about TITAN?

Titan after CASSINI and HUYGHENS

Methane-based methabolism?

Future perspectives

What do we knew about Titan?



Christiaan Huyghens discovered Titan in 1655



Typical ground-based pictures of Titan in the pre-CASSINI era.





the bearing Radar Bay Fields and **Particles Pallet** Radio/Plasma Wave Subsystem Antenna (1 of 3) Huygens Titan Probe Remote Sensing Pallet Radioisotope Thermoelectric Generator (1 of 3) 445 N Engine (1 of 2)

CASSINI spacecraft

EXOATMO Nice, Oct 2015

Titan as seen by CASSINI

Huyghen's probe (~1.30 m across)



2005.01.14

HUYGHENS result #1



 $GCMS \rightarrow$ surface wet with CH_4 that evaporated after being heated by the warmer probe.

HUYGHENS result # 2



 $GCMS \rightarrow$ Surface rich in organic compounds (ethane, cyanogen...) = complex chemistry

¹²C/¹³C measured in $CH_4 \rightarrow continuous/periodic replenishment of <math>CH_4 \rightarrow no evidences of active bio systems$



Huyghen's Descent Imager/Spectral Radiometer



What do we knew/know about Titan ?

Saturn's largest moon and 2nd biggest one in the solar System (> than Mercury)

Mass ~ 0.024 M_{earth} Radius ~0.4 R_{Earth} Mean Density ~1.9 g/cm³ (Earth's ~5.5 g/cm³) Orbital period: 16 days inclination 0.3 ° \rightarrow Saturn's equator

Eccentricity: 0.03 Tidally locked

The only moon in the solar system with a dense atmosphere

P_{Titan} ~1.45 P_{Earth}

Titan's ID after CASSINI and Huyghens

- 113 close flybys (as for 24 Sep 2015):
- Atmosphere: 98% N_2 2% CH_4 + hydrocarbons (traces)
- \$ surface temperature -180 °C (liquid methane)
- surface humidity 50% 5% methane abundance
- Iakes of ethane (76%), methane (10%) and propane (8%)
- \diamond cryovolcanoes expelling H₂O ice (?)
- \bigotimes_{\circ} subsurface mixtures of liquid amonia & water ?
- subsurface ocean





Permanent liquid hydrocarbon lakes (near north pole)



Ligeia mare (NASA - GSFC)

Weird life on the surface of Titan ?

Methanogenic archea could live on Titan using a number of

metabolic pathways, where terrestrial O_2 is changed by H_2

Such organisms would inhale H_2 in place of O_2 , metabolize it with

acetylene instead of glucose, and exhale methane instead of carbon dioxide

(e.g. Abbas & Schulze-Makuch 2002; McKay & Smith 2005)

Acetylene \rightarrow C₂H₂ + 3H₂ \rightarrow 2CH₄

Acetate \rightarrow CH₃COO- + H+ \rightarrow CH₄ + CO₂

 \rightarrow Methanogenic archea on Earth can survive on the energy levels released by these Fearthons^{Oct 2015}

BUT:

If microorganisms are consuming Hydrogen and Acetylene on Titan's surface, the abundances of them should be measurably lower than otherwise expected (McKay & Smith 2005)

- Indeed, Strobel (2010) found a greater abundance of H_2 in the upper atmospheric layers of Titan compared to the lower layers and
- Clark et al. (2010) fail to find Acetylene on the surface of Titan from CASSINI's VIMS data. Other explanations are possible \rightarrow \rightarrow
- → → Must wait for more flybys EXOATMO Nice, Oct 2015

February 2015:

Stevenson et al., Sci Adv 2015, 1: →

As liquid bylayer membrane NOT possible without liquid water.

ALTERNATIVE: Molecular simulations ->

New type of membrane, composed of small organic nitrogen compounds, capable of forming and functioning in liquid methane at cryogenic temperatures.

The "azotosome" has properties similar to lipidic membranes. It can be formed from compounds observed in Titan's atmosphere.



Acrylonitrile azotosomes and final vesicule

(CH2CHCN)

THANK YOU!